



Variability of Space Weather over Africa from IHY/ISWI Observational Facilities between 2006 and 2012

Babatunde Rabiou

National Space Research & Development Agency, NASRDA, Abuja, Nigeria

*Space Physics Laboratory, Physics dept., Federal University of
Technology, Akure, Nigeria*

Email: tunderabiou@yahoo.com, babatunde.rabiou@nasrda.gov.ng



outline

- Africa
- Gains of UNBSSI/IHY/ISWI
- Infrastructural development
- Scientific results
- Conclusions



Africa !

- A continent
- 54 individual nations
- Multi-lingual structure
- English, French, Portuguese, Arabic, Spanish
- ~ 30 billion km²
- ~ 850 million people
- ~14% of the World population





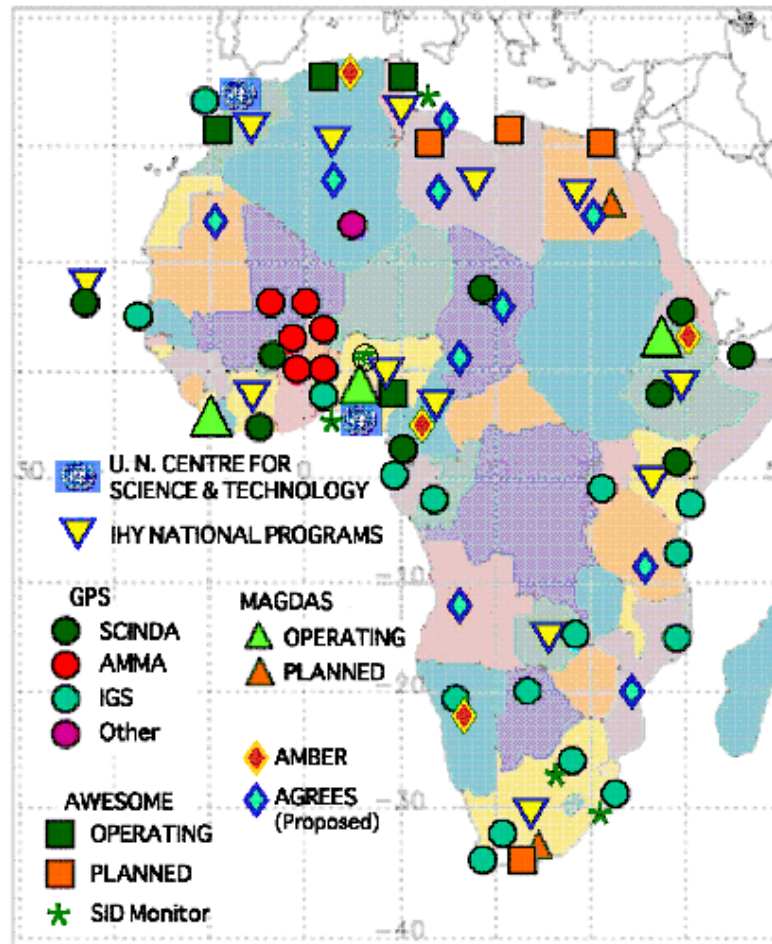
IHY/ISWI facilities in Africa

- Over 17 magnetometers (MAGDAS and AMBER)
- more than 25 GPS receivers (SCINDA and others)
- well over 50 ionospheric RF sounders (Ionosonde, SID monitor and AWESOME)



Signatures of UN and IHY in Africa

– strictly after Barbara Thompson 2007





Gains of IHY/ISWI in Africa

- Knowledge & technological transfer
- Positive collaboration
- Availability of Research facilities for internationally competitive research.
- Publication of scholarly articles
- Windows of postgraduate opportunities
- Control of brain drain
- Development of Research in Basic Space Science
- Capacity building
- Bridge between North & South
- strong intra–continental partnerships amongst African scientists

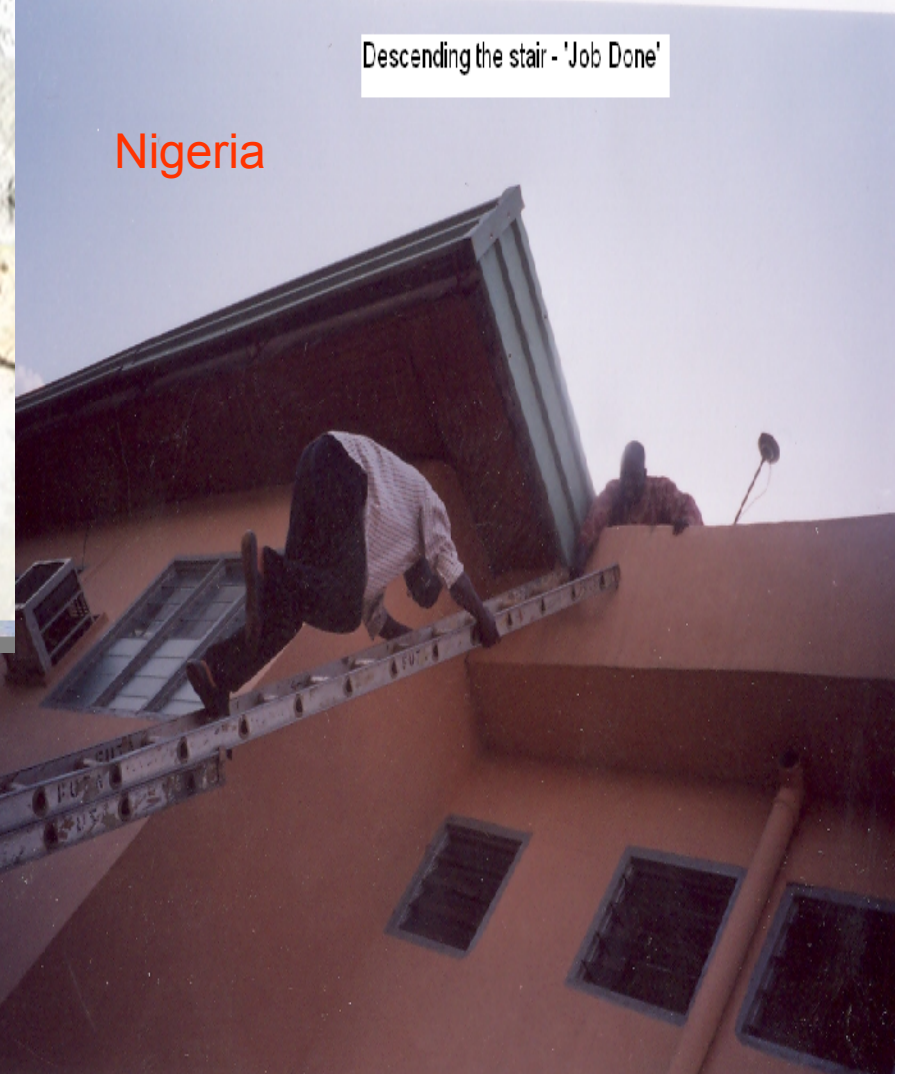




MAGDAS at ILORIN, Nigeria. August 2006
UN/Ecuador Workshop on the Interspace Weather Initiative 9-12 October 2012 Quito, Ecuador



Climbing the stairs





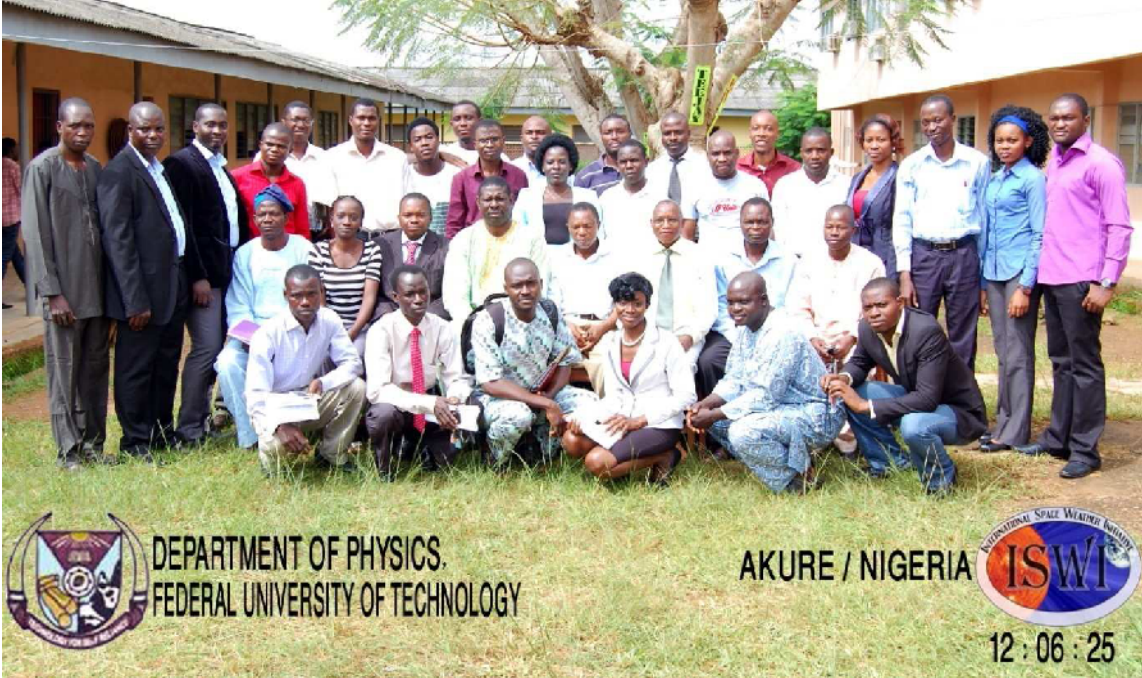
Summer Schools

- 41 African graduate students & Postdocs
- From 14 African countries
- 15 Instructors



CYRIL ONWUMECHILI SCHOOL ON PHYSICS OF GEOMAGNETIC PHENOMENA

24 - 30 June 2012,
Akure, Nigeria



DEPARTMENT OF PHYSICS
FEDERAL UNIVERSITY OF TECHNOLOGY

AKURE / NIGERIA



UN/Ecuador Workshop on the Int'l S



or



Some Results



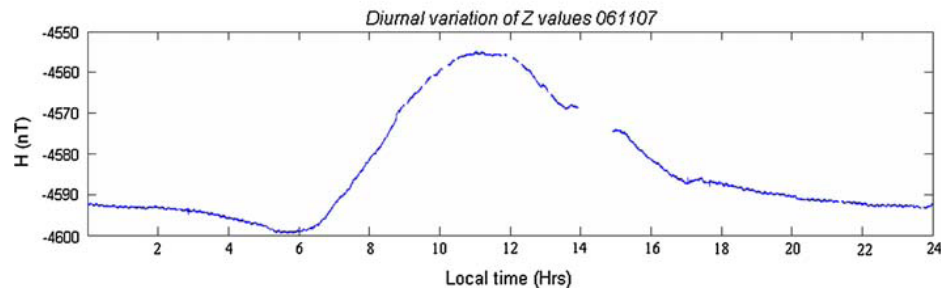
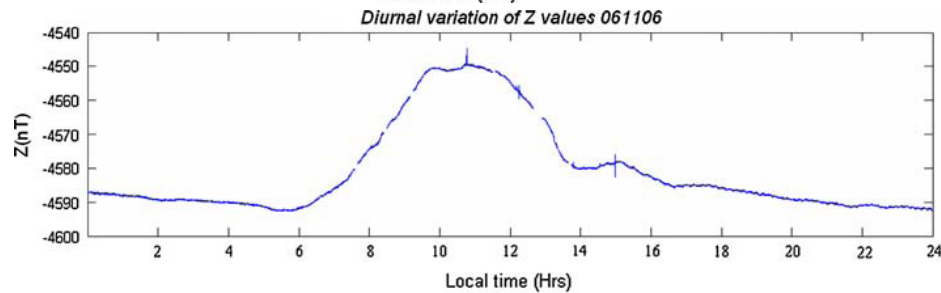
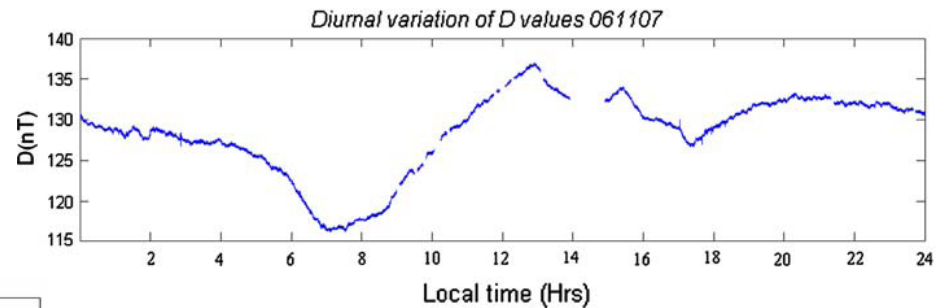
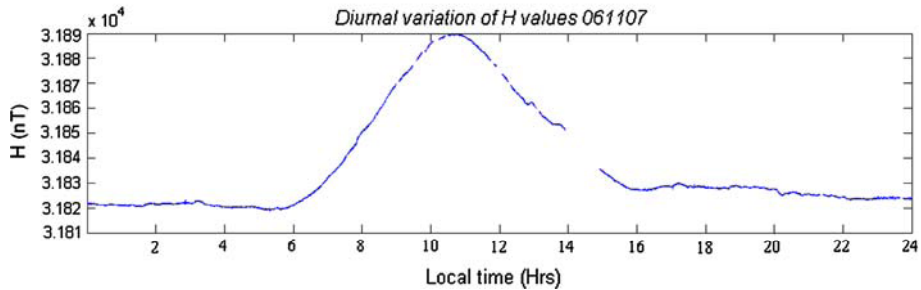
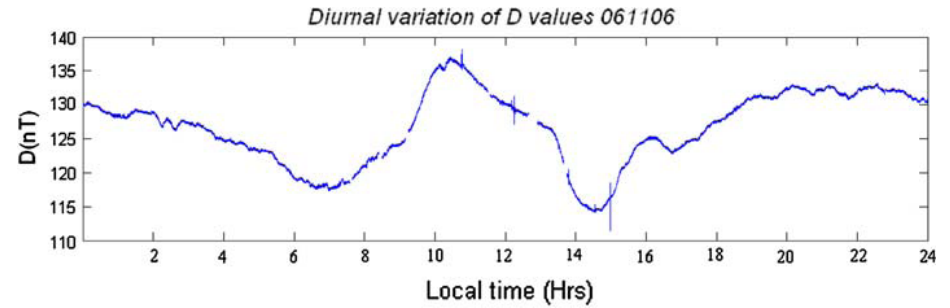
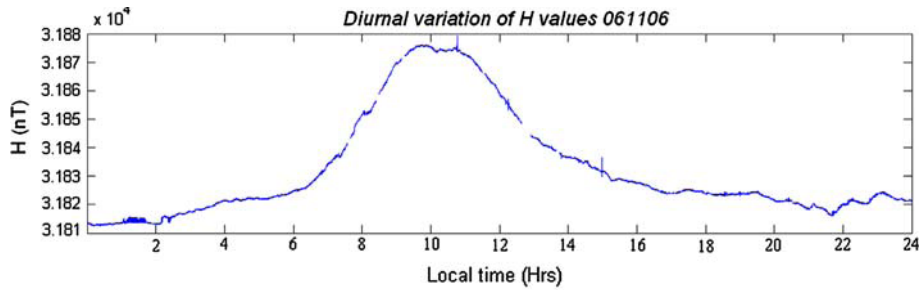
Earth Moon Planet (2009) 104:173–179
DOI 10.1007/s11038-008-9290-7

Preliminary Results from the Magnetic Field Measurements Using MAGDAS at Ilorin, Nigeria

**A. B. Rabiū · I. A. Adimula · K. Yumoto · J. O. Adeniyi ·
G. Maeda · MAGDAS/CPMN Project group**



Diurnal variation of H D Z on 6 & 7 Nov 2006



- Diurnal variations
- Day-to-day variability of ionospheric process etc
- Rabiou et al 2009



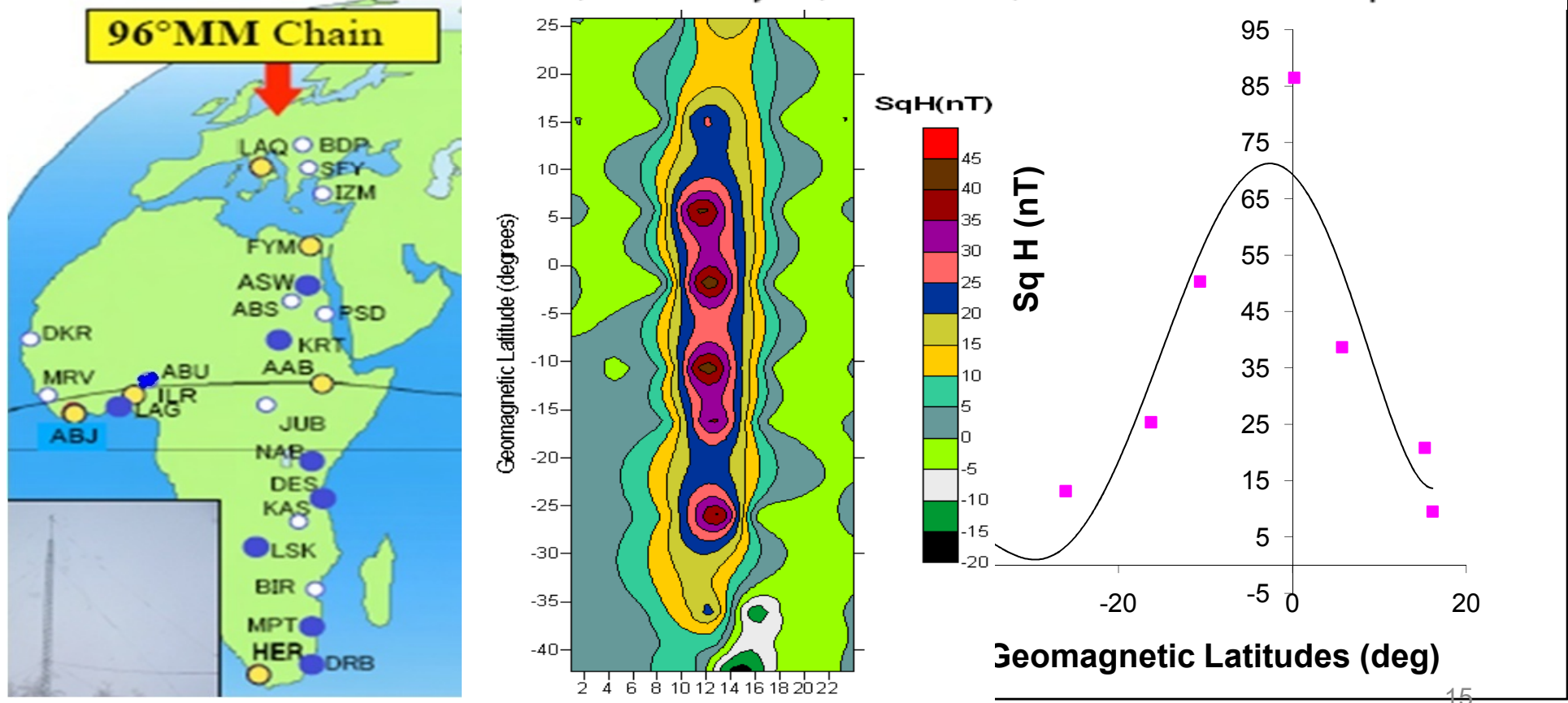
A new index to monitor temporal and long-term variations of the equatorial electrojet by MAGDAS/CPMN real-time data: *EE*-Index

T. Uozumi¹, K. Yumoto¹, K. Kitamura², S. Abe¹, Y. Kakinami¹, M. Shinohara¹, A. Yoshikawa¹, H. Kawano¹,
T. Ueno³, T. Tokunaga³, D. McNamara⁴, J. K. Ishituka⁵, S. L. G. Dutra⁶, B. Damtie⁷,
V. Doumbia^{*8}, O. Obrou⁸, A. B. Rabiou⁹, I. A. Adimula¹⁰, M. Othman¹¹,
M. Fairos¹¹, R. E. S. Otadoy¹², and MAGDAS Group¹

This index provides information that should clarify the situation of solar-geospace coupling and atmosphere-ionosphere coupling in the magnetic equatorial region

Ionosphere over Africa: Results from Geomagnetic Field Measurements During International Heliophysical Year IHY

A.B. Rabiou^{1,2}, K.Yumoto³, E.O. Falayi^{2,4}, O.R.Bello², MAGDAS/CPMN Group³





Seasonal variation of Sq(H) along the latitudes

- Sq (H) is greater in all seasons in the neighbourhood of dip equator
- Obviously due to EEJ effect
- Max effect at Autumn (Sept) Equinox

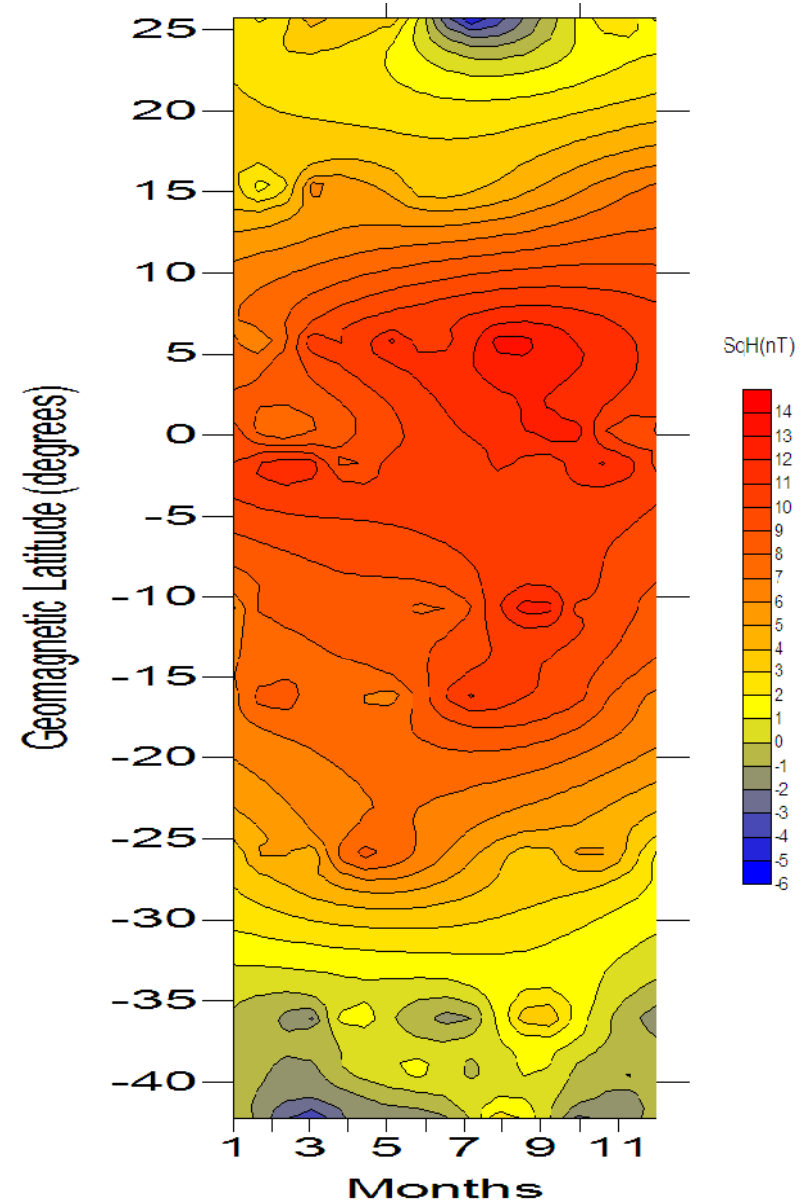


Figure : Seasonal Variation of SqH (nT)

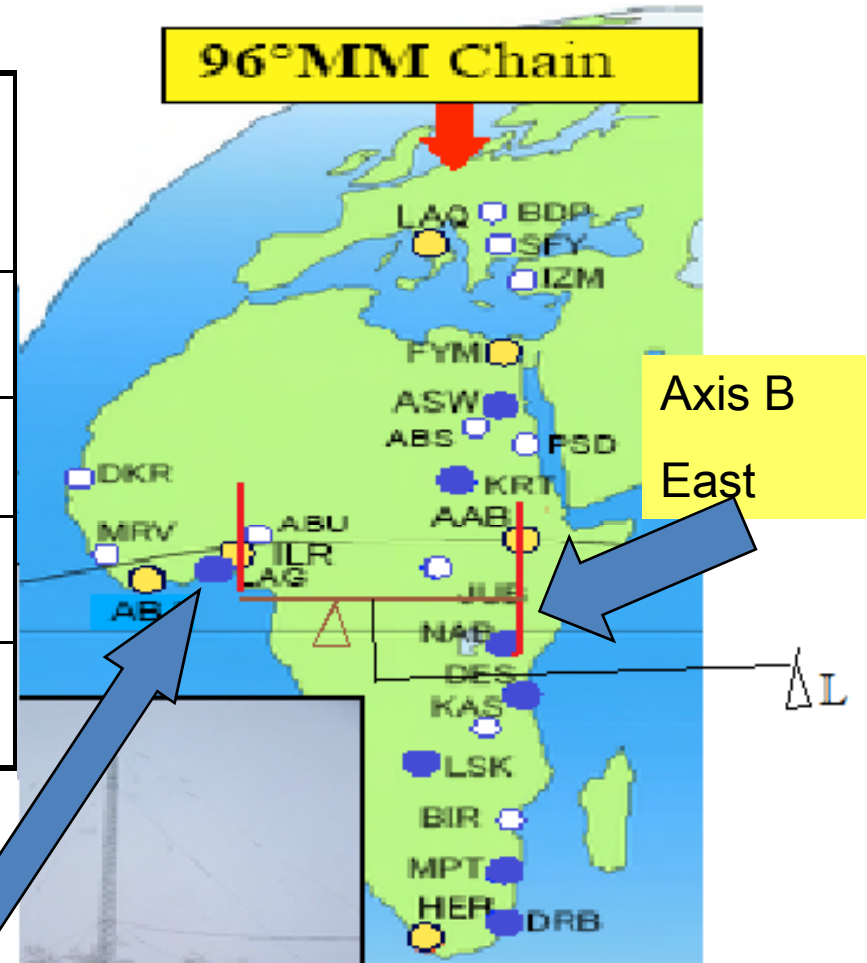


EEJ in Africa

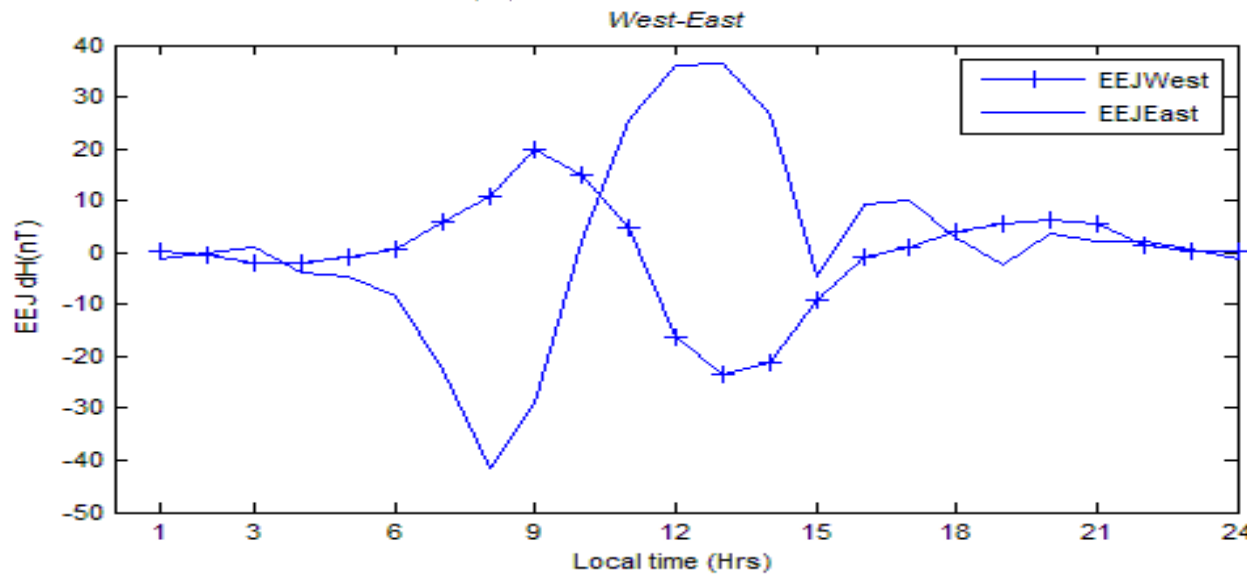
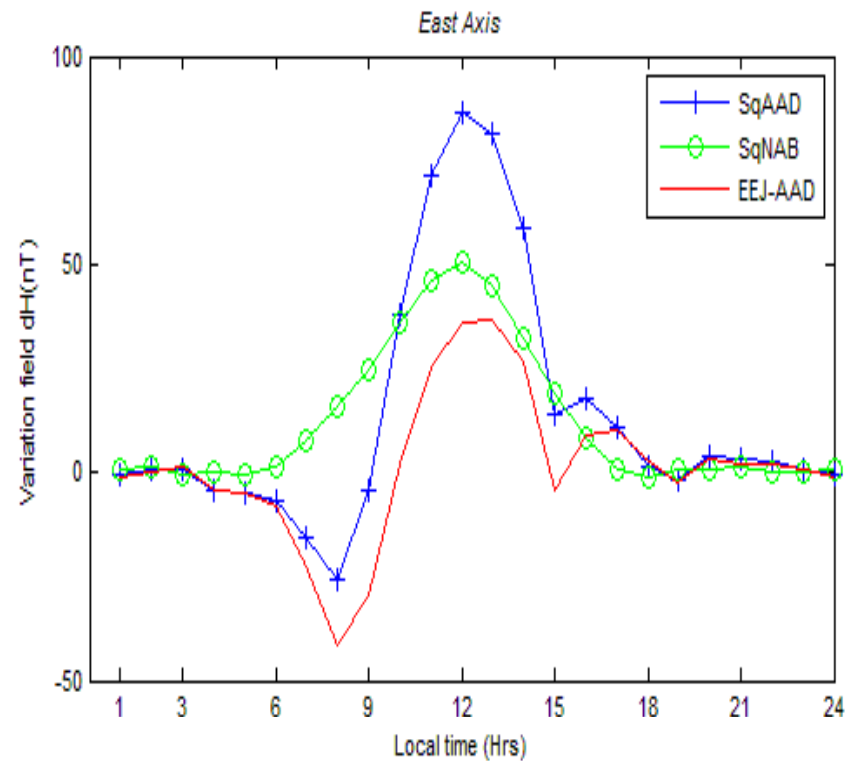
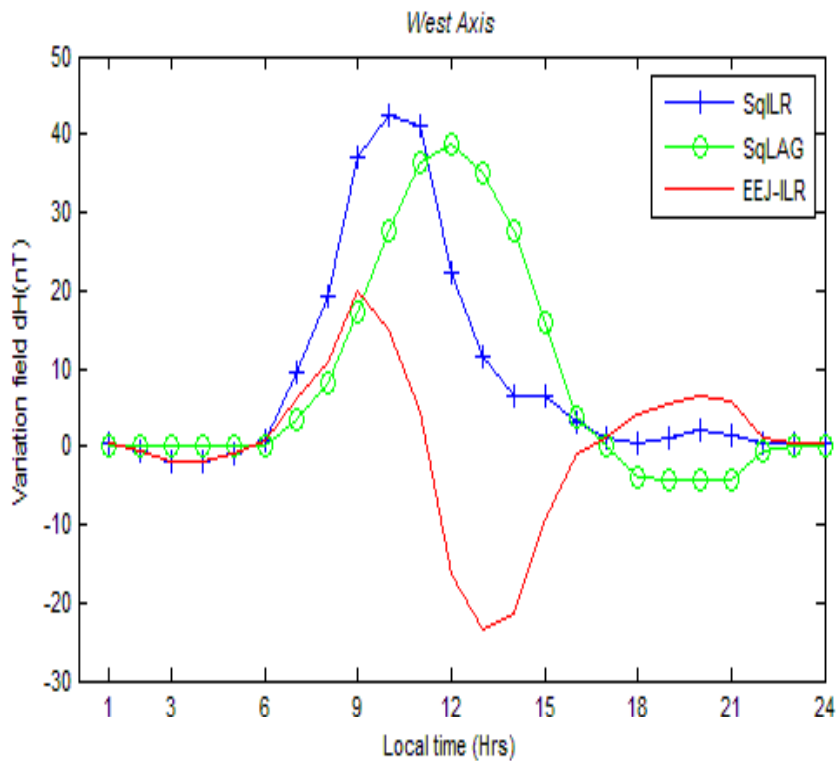


Coordinates of the Stations

OBS	GMLat°	GLong ° E	GLat°
ILR	-1.82	4.67	8.50°N
LAG		3.43	3.42°N
AAB	0.18	38.77	9.04°N
NAB		36.80	1.16°S



Separation of axes, $\Delta L = 33.735^\circ = 3744.585 \text{ km}$

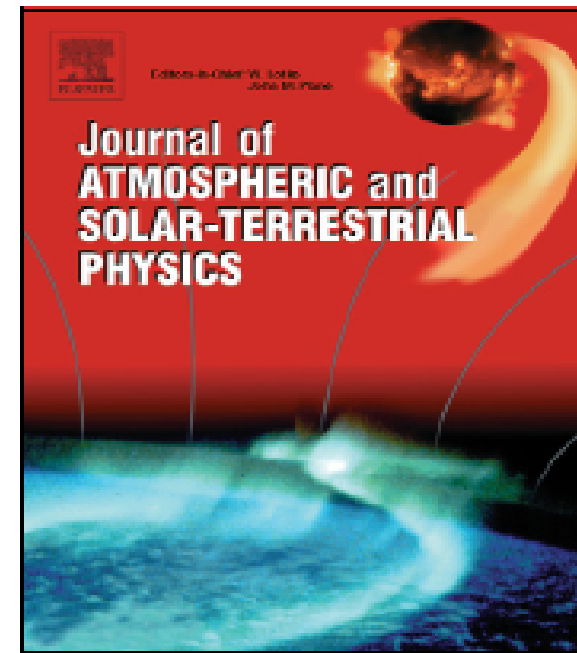


Western EEJ appears weaker than Eastern EEJ!
 It is as if there is a process of re-injection of energy as Jet flows eastward

Author's Accepted Manuscript

Climatology of the inter-hemispheric field-aligned currents system over the nigeria ionosphere

O.S. Bolaji, A.B. Rabiou, E.O. Oyeyemi, K. Yumoto



www.elsevier.com/locate/jastp

PII: S1364-6826(12)00184-8
DOI: <http://dx.doi.org/10.1016/j.jastp.2012.07.008>
Reference: ATP3656

To appear in: *Journal of Atmospheric and Solar-Terrestrial Physics*

Received date: 19 March 2012
Revised date: 23 June 2012
Accepted date: 27 July 2012

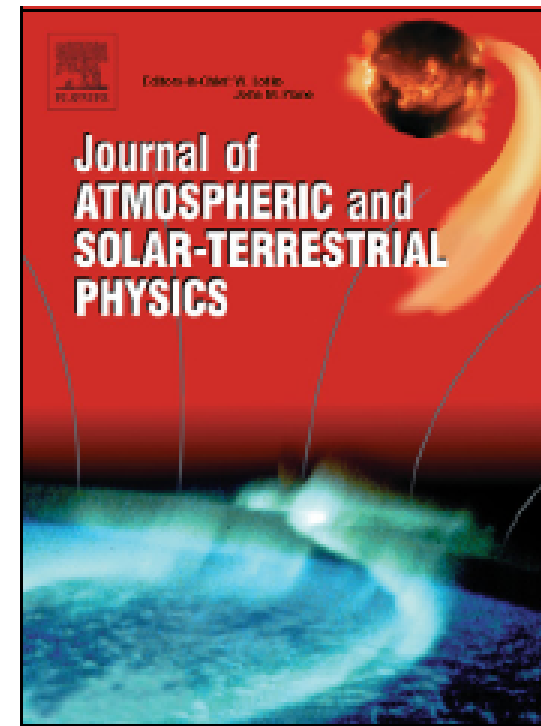
Cite this article as: O.S. Bolaji, A.B. Rabiou, E.O. Oyeyemi and K. Yumoto, Climatology of the inter-hemispheric field-aligned currents system over the nigeria ionosphere, *Journal of Atmospheric and Solar-Terrestrial Physics*, <http://dx.doi.org/10.1016/j.jastp.2012.07.008>



Author's Accepted Manuscript

Climatology of the inter-hemispheric field-aligned currents system over the nigeria ionosphere

O.S. Bolaji, A.B. Rabiou, E.O. Oyeyemi, K. Yumoto



www.elsevier.com/locate/jastp

O.S. Bolaji, A.B. Rabiou, E.O. Oyeyemi and K. Yumoto, Climatology of the inter-hemispheric field-aligned currents system over the nigeria ionosphere, Journal of Atmospheric and Solar-Terrestrial Physics, <http://dx.doi.org/10.1016/j.jastp.2012.07.008>



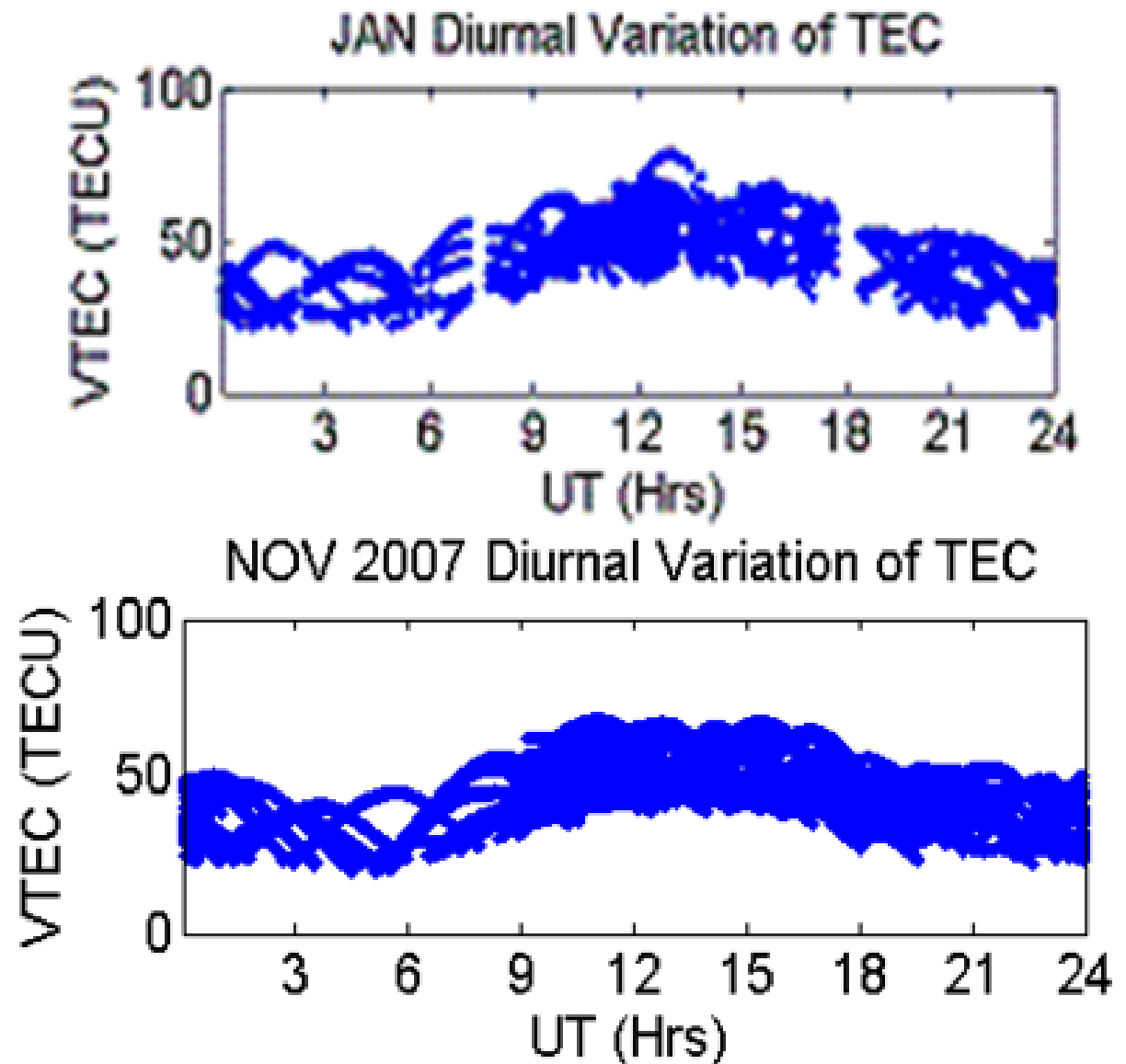
IHFACs Climatology

- The IHFACs magnetic field variation flow in opposite direction of the winter northern-hemisphere.
- Dusk-side IHFACs was confirmed & are weakly northbound in all the seasons.
- Diurnal, monthly mean and seasonal variations of IHFACs exist and exhibit downward & upward inter-hemispheric field-aligned sheet current that appears as a pair at all local times.
- IHFACs exhibit longitudinal variability



TEC STUDIES

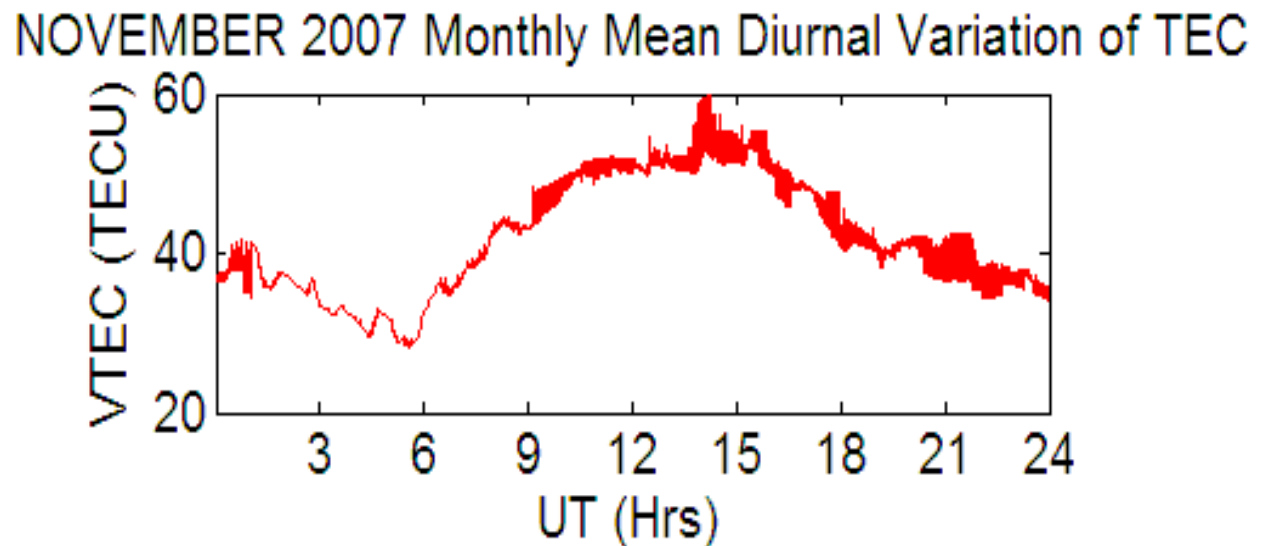
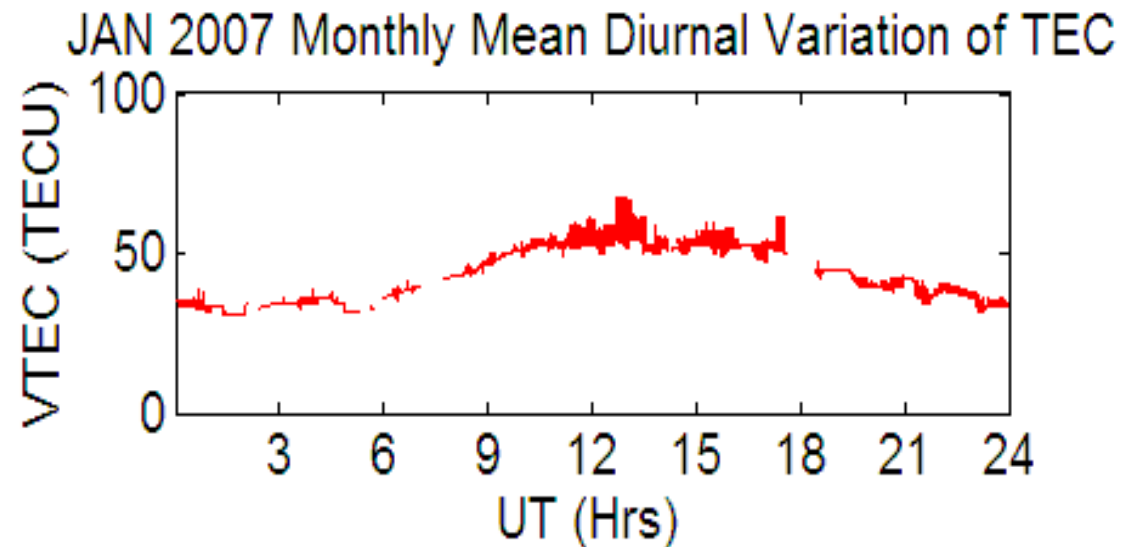
Mass plots
of the
Diurnal
Variation of
VTEC as
observed
from the
data from all
the visible
PRN
over Akure



Diurnal Variation of VTEC over Akure

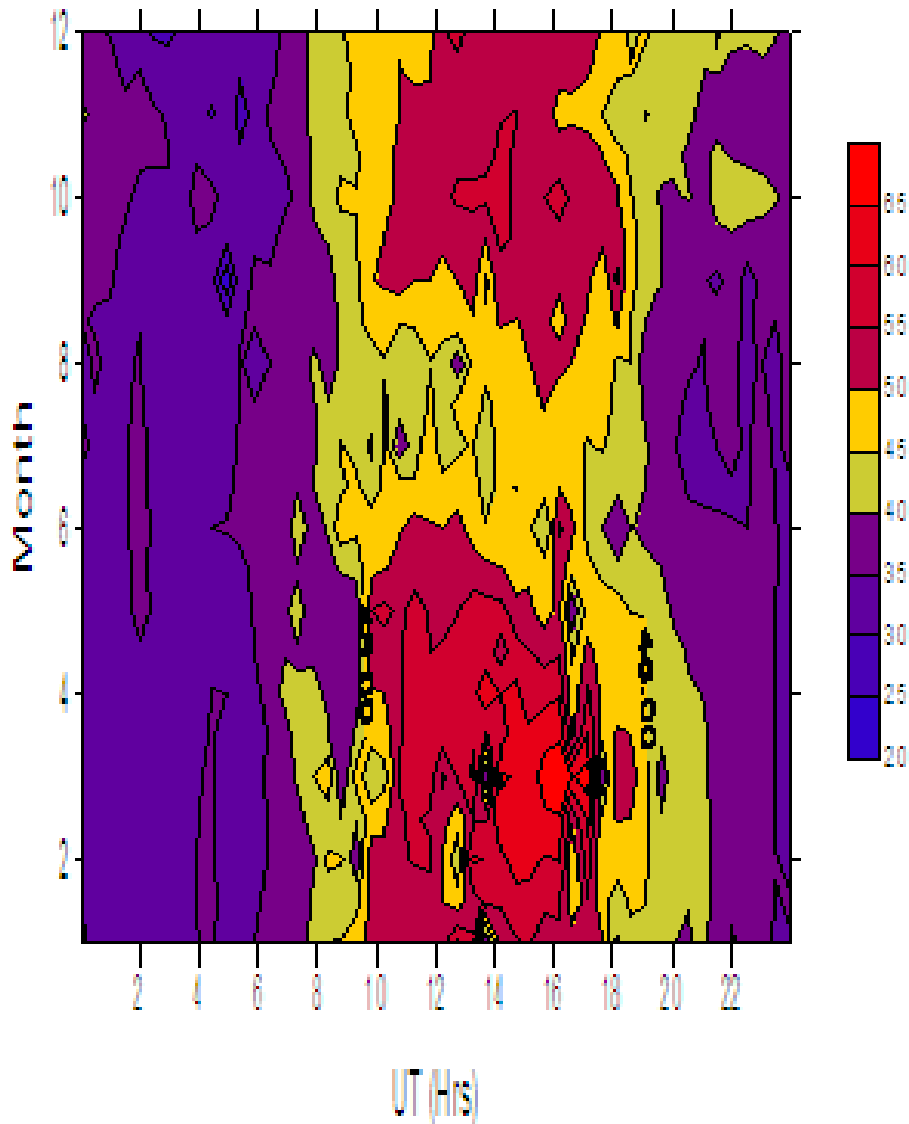
✓ pre-dawn minimum
for a short period of
time followed by steep
early morning
increase.

✓ TEC reaches
maximum value
between 1300UT
(1400LT) & 1400UT
(1500LT)





Annual VTEC variation at Akure, Nigeria

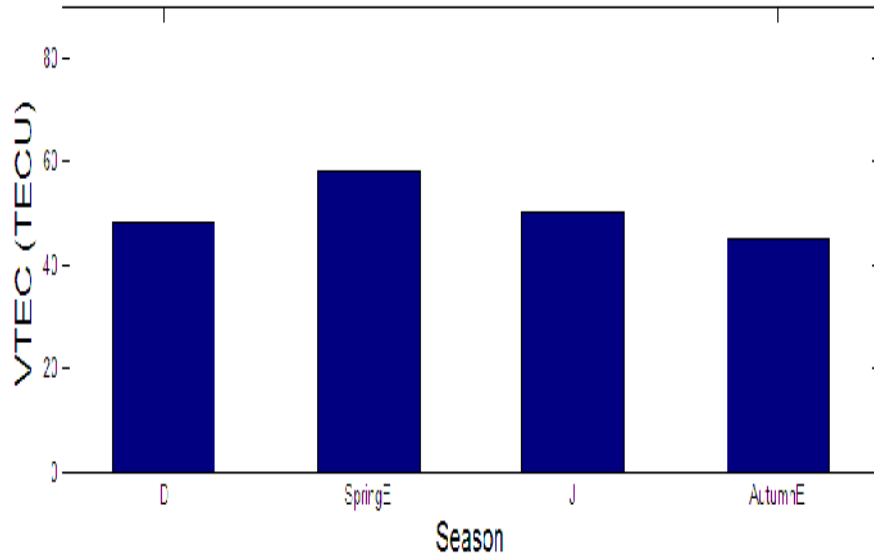


- pre-dawn minimum for a short period of time followed by steep early morning increase.
- Attain maximum between 14.00UT and 16.00UT.
- maximizes during Equinox months, minimizes during winter months
- The semiannual variation of TEC is asymmetry with maximum in spring Equinox

Rabiu et al 2011



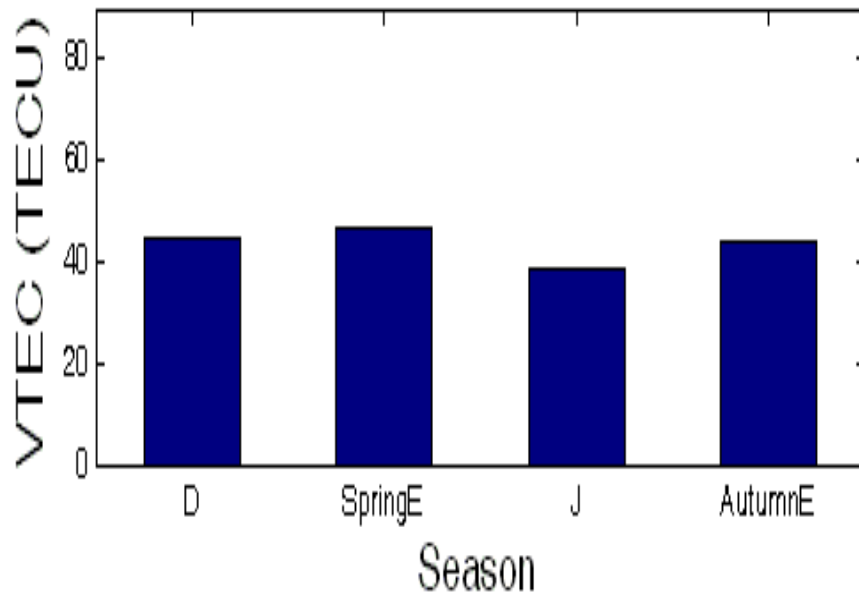
Seasonal Variation of VTEC 2007



Seasonal variation of TEC

- The semiannual variation of TEC is asymmetry with maximum in spring Equinox.

Seasonal Variation of VTEC 2008

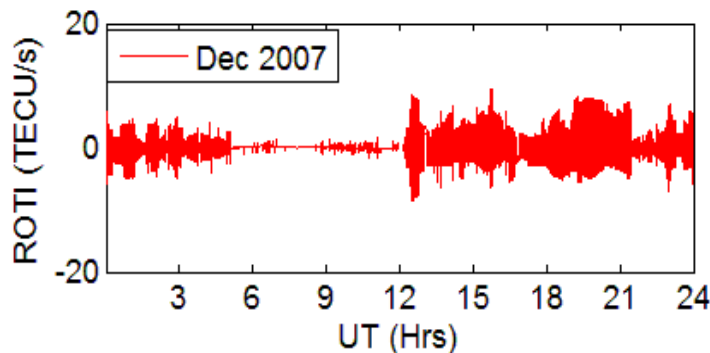
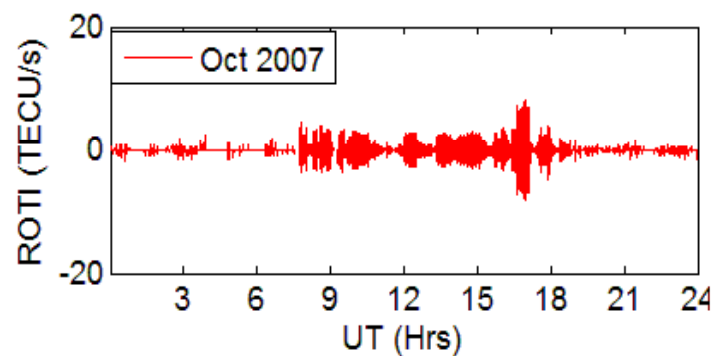
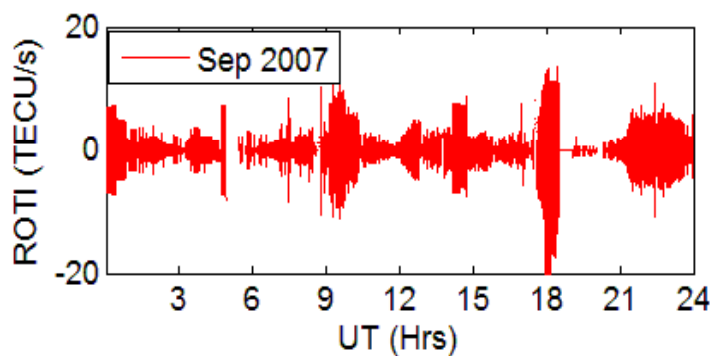
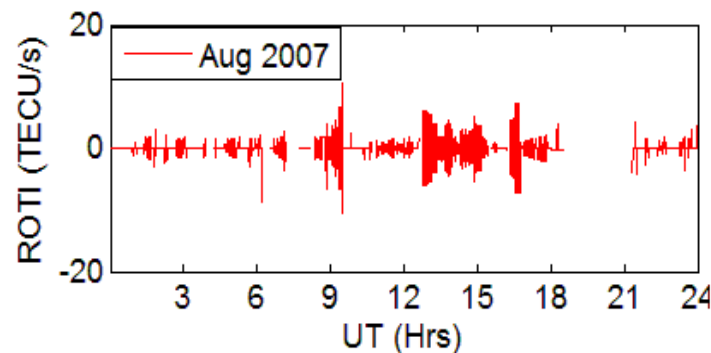
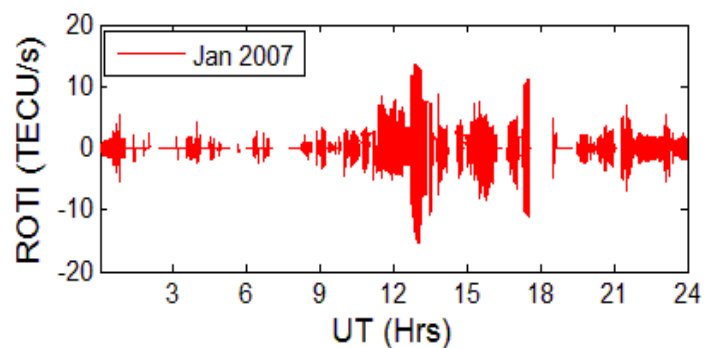




ROT

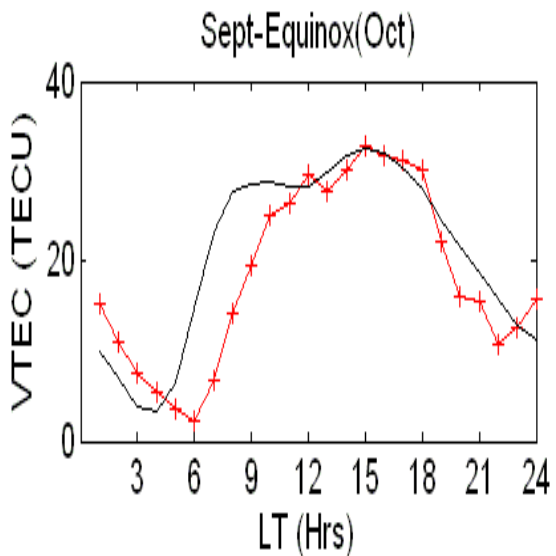
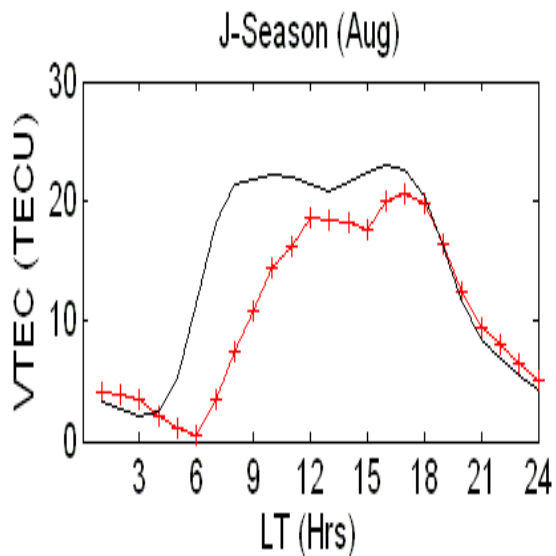
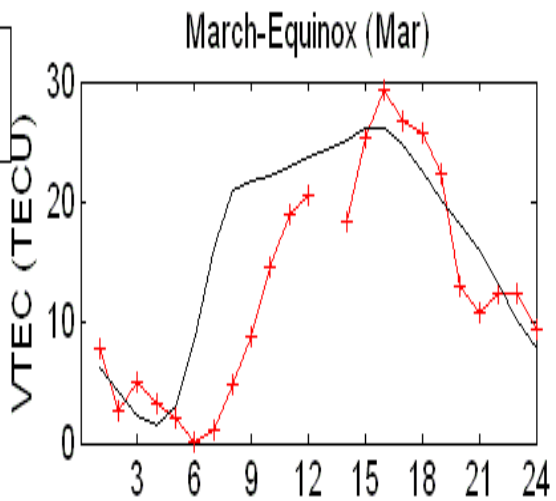
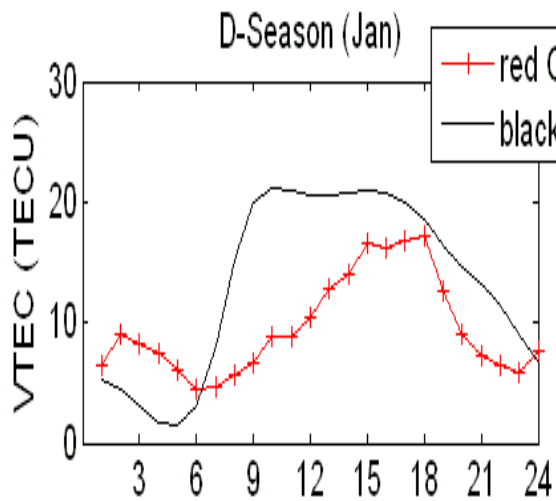
✓ More TEC
Fluctuations
in daytime.

✓ Seems to
lack
seasonal
dependent



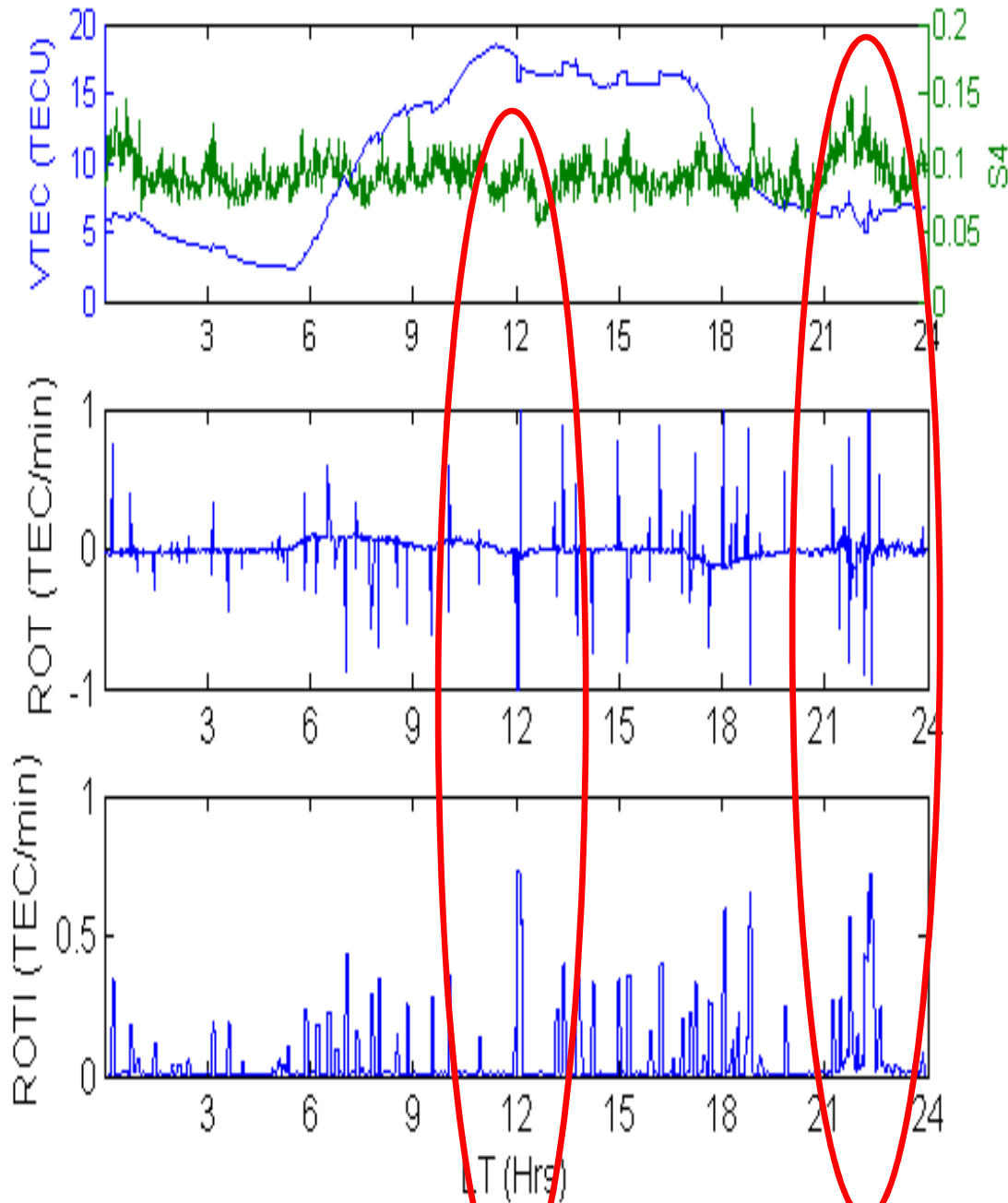
$$\text{ROT} = [\text{TEC}(t) - \text{TEC}(t+1)] / 1$$

Unit TECU/sec



IRI Under- & over- estimate the values of TEC at different times in all the seasons considered.

IRI & Observed TEC @ Akure 2010

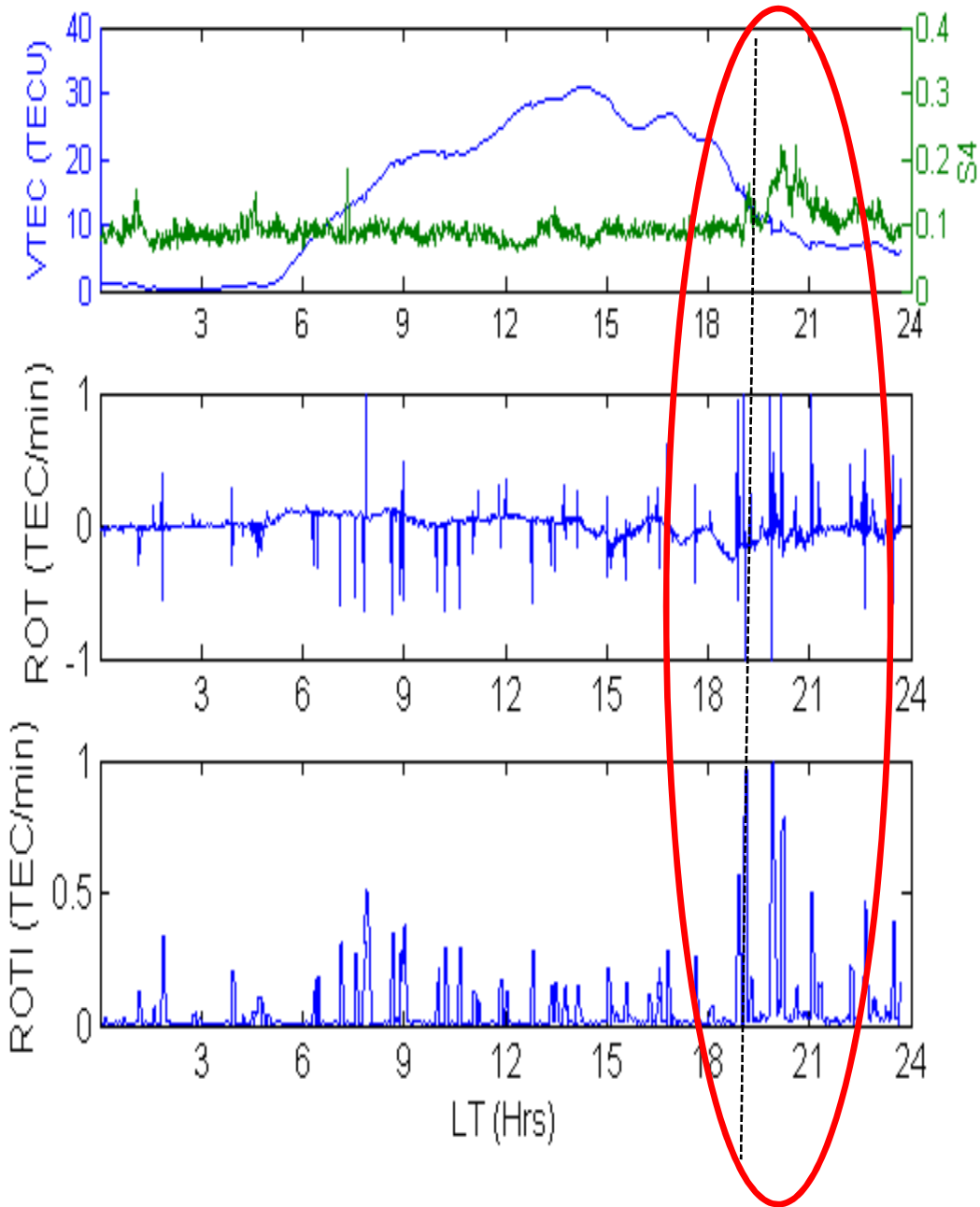


VTEC, S4, ROT & ROTI,
17th Jan 2010
Ap = 1

S4, ROT and ROTI give
information on the di
mension
of irregularities

S4 is sensitive to scale
< the Fresnel scale

ROTI > 0.5 corresponds
to scale lengths of few
km



VTEC, S4, ROT &
ROTI,

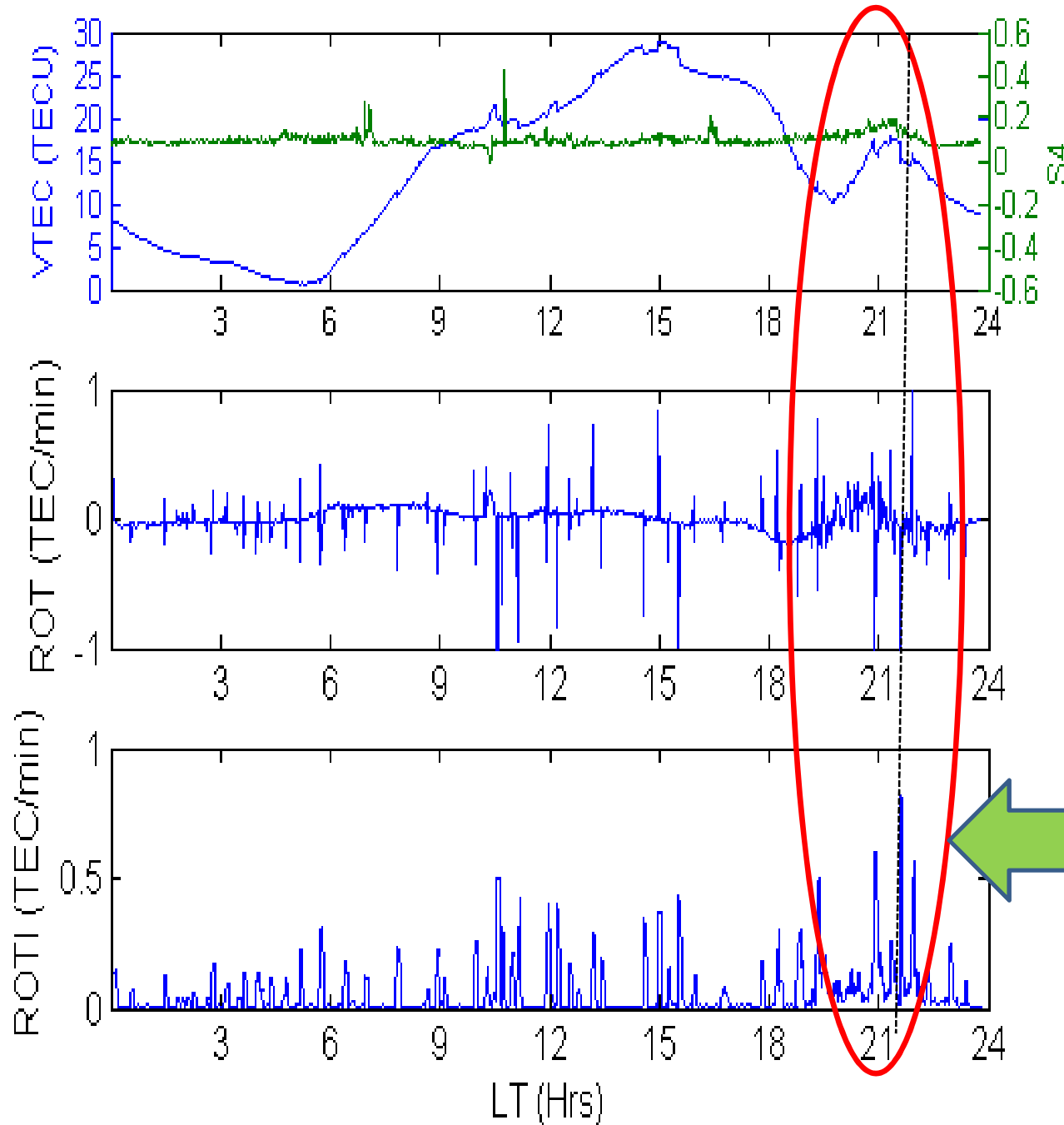
27th August 2010

$A_p = 14$

ROTI :
standard
deviation of
ROT at 5 mins
interval

VTEC, S4, ROT &
ROTI,
3rd March 2010

Ap = 8



ROTI gave a more pronounced representation of TEC fluctuation than S4



Equivalent ranges of TEC & Dst at Akure, Nigeria. April 1-15, 2010

Activity Level	Dst (nT)	Mean TEC (TECU)
Low	Dst > -20	16.713
Medium	-20 > Dst > -50	16.851
High	-50 > Dst > -100	20.138

TEC increases with increasing magnetic activity.

Measured TEC could serve as proxy for monitoring ionospheric responses to magnetic activity

Rabiu & Abdulrahim (2012): PAPER ID: IAC-12, D5,3,6,x12869 , IAF, Naples, Italy

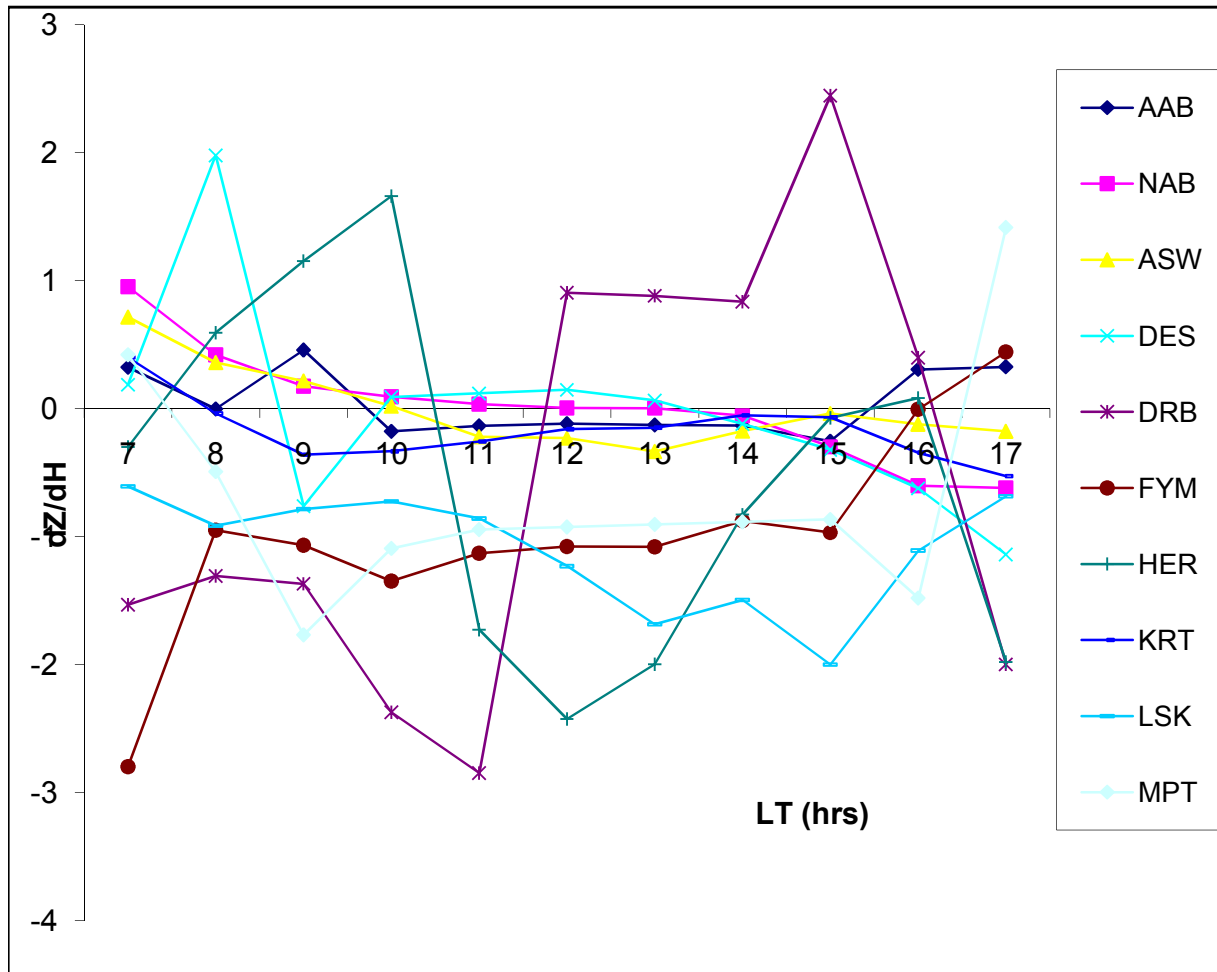


A STUDY OF ELECTROMAGNETIC INDUCTION DUE TO GEOMAGNETIC EFFECT AT AFRICAN 96° MAGNETIC MERIDIAN

Rabiu, Yumoto, Yamazaki, Cardinal, MAGDAS/CPMN Group (2009)

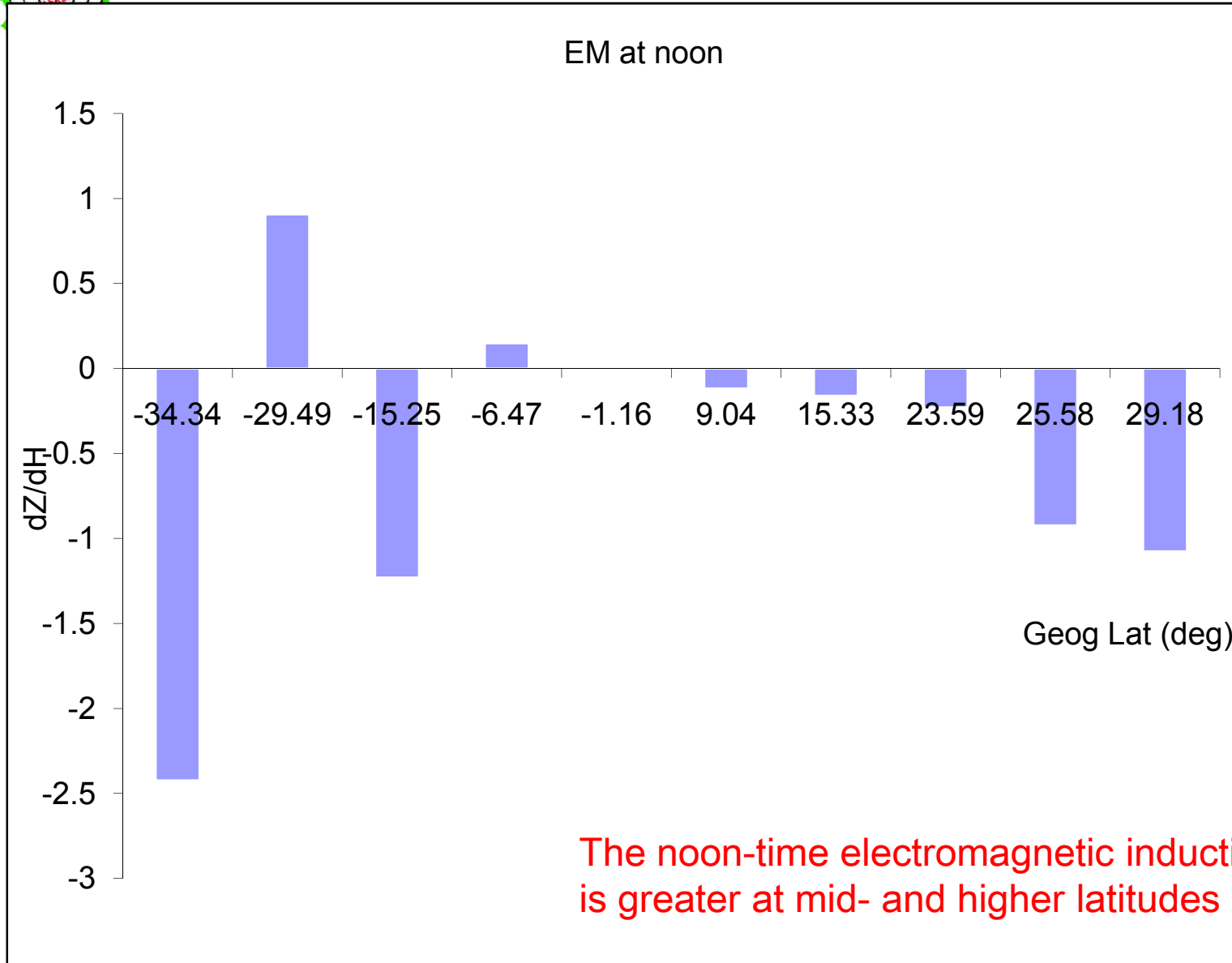


Daytime variation of dZ/dH at different stations along 96°MM



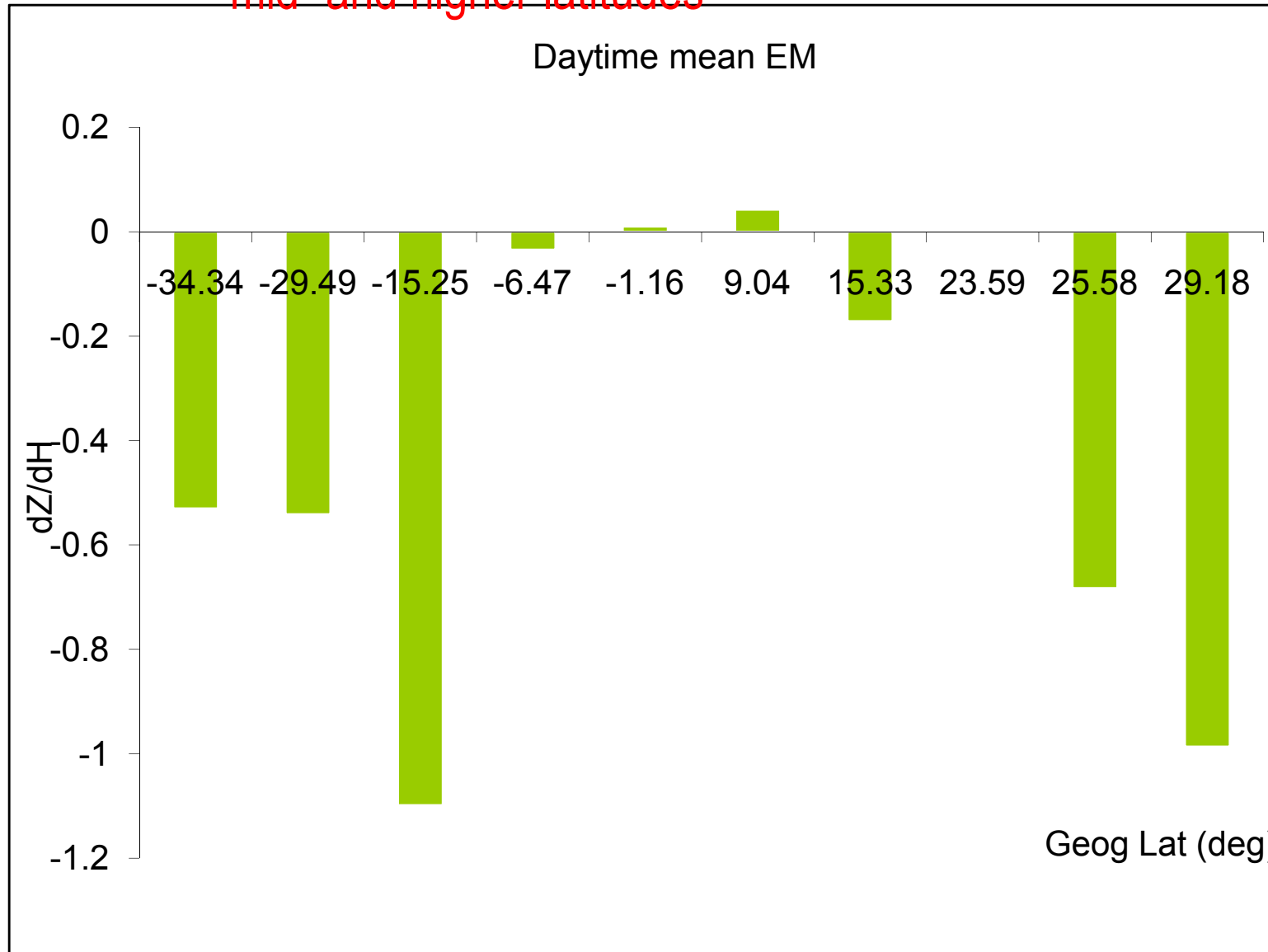
- ✓ Stations outside EEJ tends to have more EM induction effects
- ✓ The EM induction is minimal at all stations at about local noon
- ✓ More effect at sunrise and sunset periods
- ✓ Return currents could explain this observation. (Rabiu & Nandini, 2008)

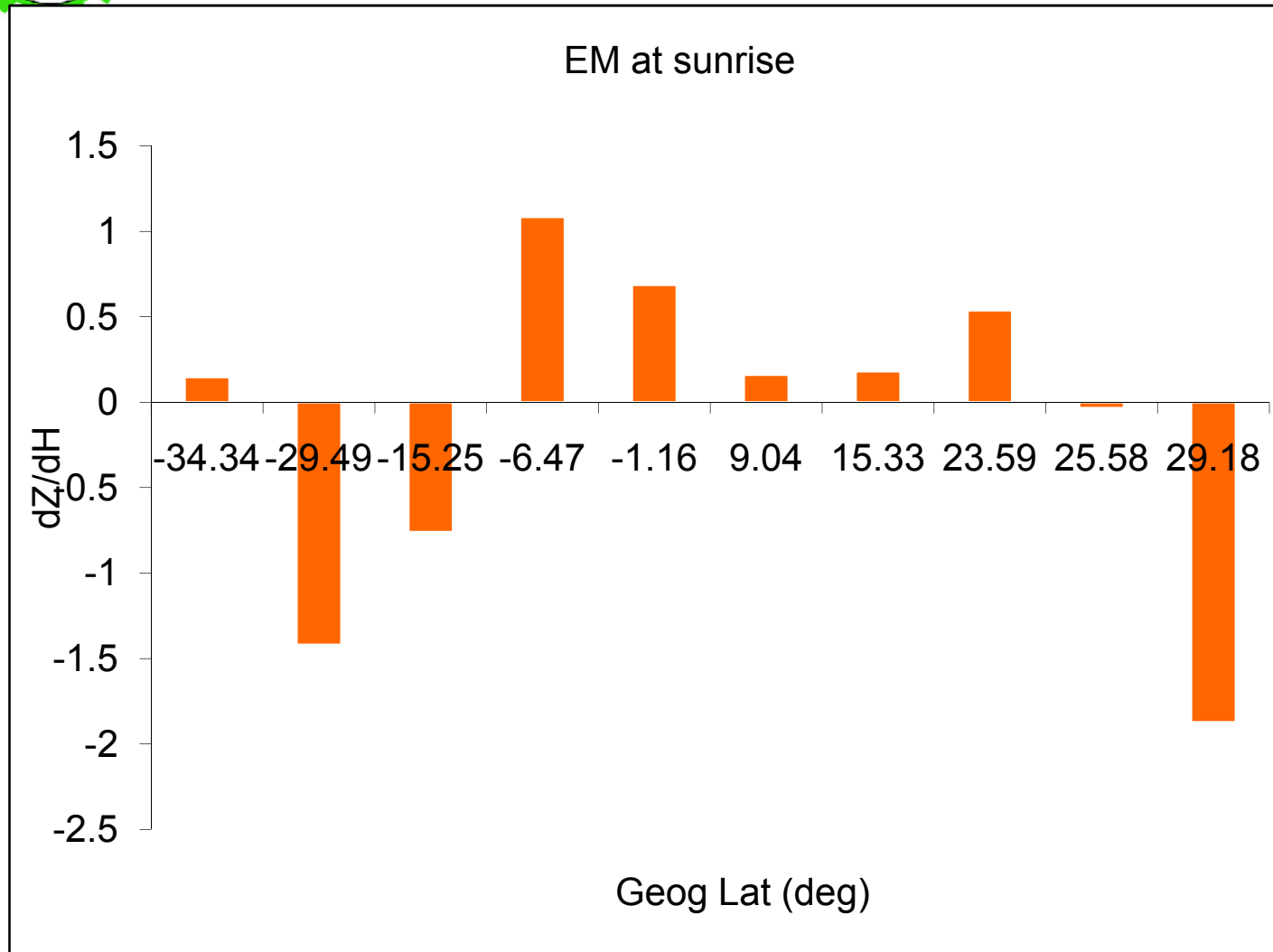
The electromagnetic inductive response is negligible around local noon, when the electrojet source field has zonal symmetry, in all sectors



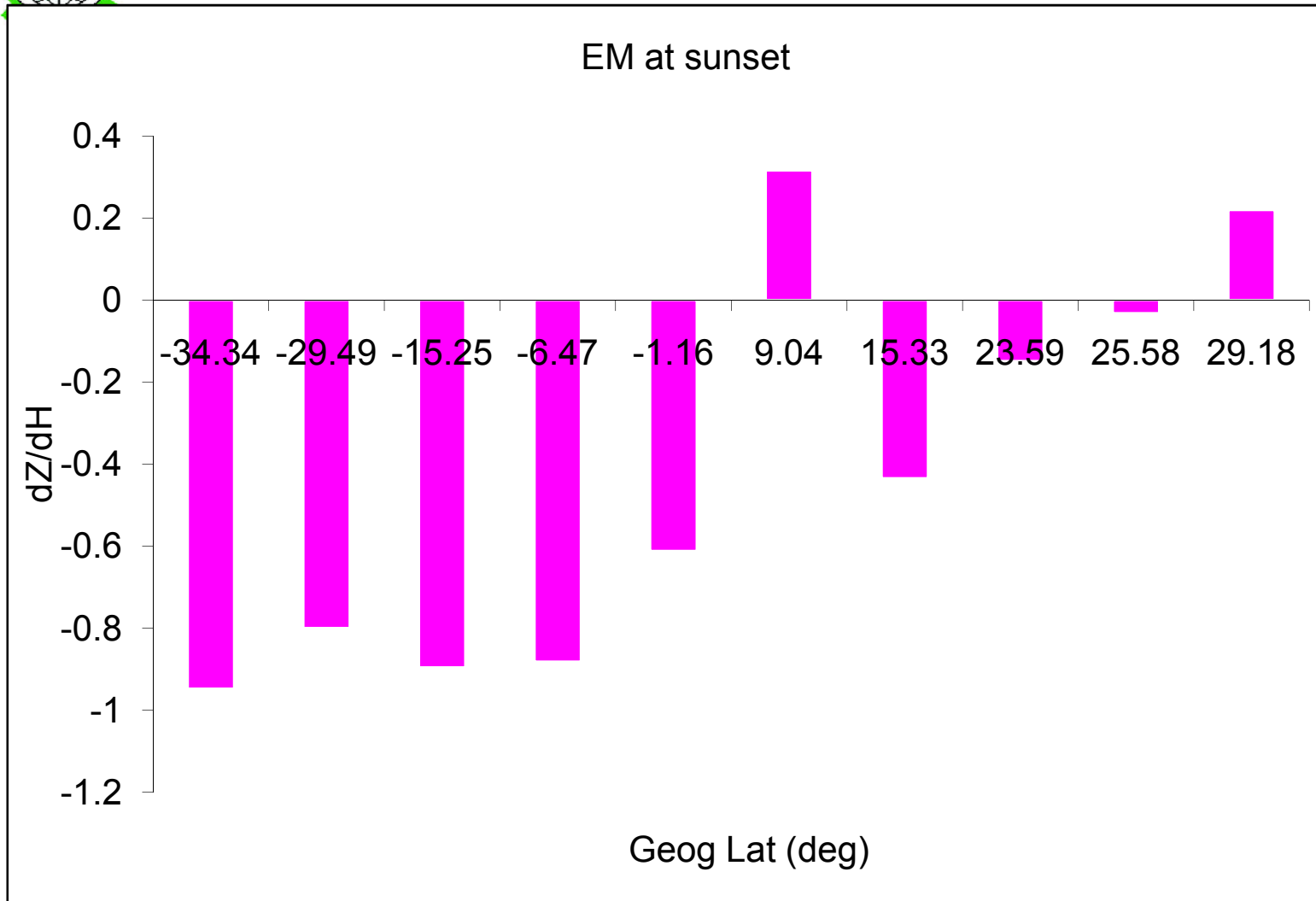


The day-time electromagnetic inductive response is greater at mid- and higher latitudes





The sunrise electromagnetic inductive response is greater at mid- and higher latitudes



The sunset electromagnetic inductive response is greater at southern hemisphere?



Conclusions

- ❑ We highlighted the contributions of UN BSSI (1992-2006), International Heliophysical Year IHY (2007-2009) and International Space Weather Initiative ISWI (2009-2012) to Africa
- ❑ Over 17 magnetometers (MAGDAS and AMBER), more than 25 GPS receivers (SCINDA and others), and well over 50 ionospheric RF sounders (Ionosonde, SID monitor and AWESOME) are now operational in Africa
- ❑ Some scientific results have been presented



Thank You