SWx Data Center for Asia Region

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Korea Astronomy & Space Science Institute

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Young-Deuk PARK, Ph.D

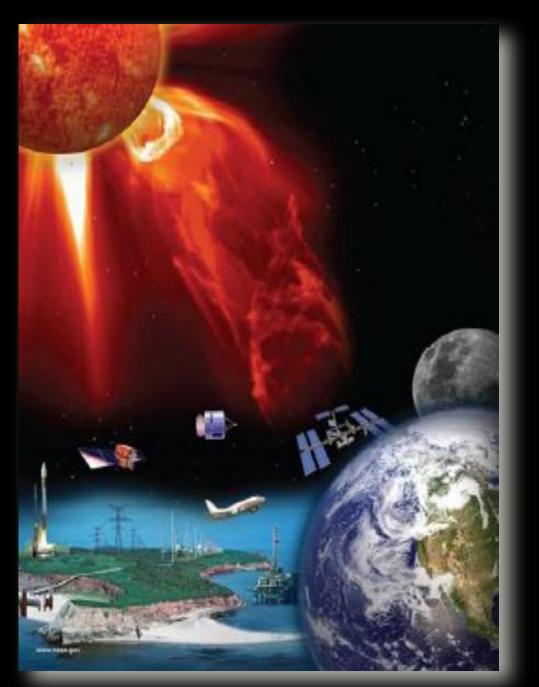
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Introduction

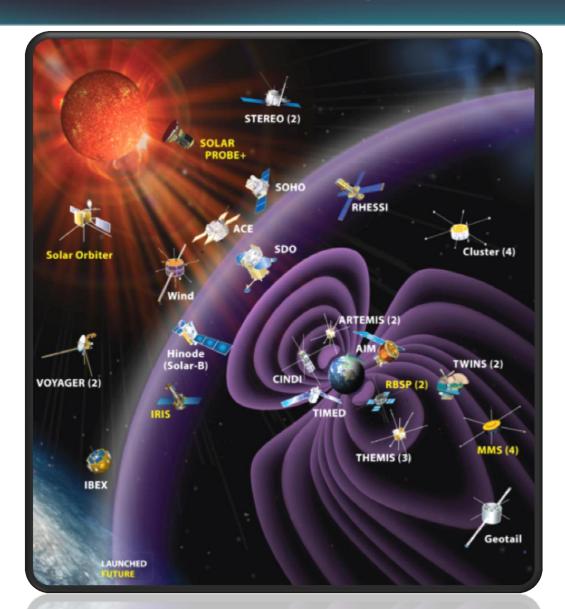


21 century is space era and space weather is important as much as weather on the ground human life.

Solar activity is the main source of the space weather so we need to monitoring the solar activity in detail

We have to the responsibility as space scientist, predict the space weather and minimize the global disaster.

Human effort to space weather prediction



Enough Obs. Data

How to select Data for your research

How to get the Data for your research



Data Center : General

Data Center : General

- One of the main concerns is **business continuity**
 - Provide a reliable infrastructure for IT operations
- Information security
 - Secure environment which minimizes the chances of a security breach
 - Keep high standards for assuring the integrity and functionality of its hosted computer environment
- The equipment may be used to:
 - Operate and manage a carrier's telecommunication network
 - Provide data center based applications directly to the carrier's customers
 - Provide hosted applications for a third party to provide services to their customers
 - Provide a combination of these and similar data center applications

High performance DC

Standardization/consolidation:

- Reduce the number of data centers a large organization may have.
- Reduce the number of hardware, software platforms, tools and processes
- Organizations :
 - provide increased capacity and performance.
 - easier to manage.

9 Virtualize:

- Helps to lower capital and operational expenses, and energy consumption.
- Virtualization technologies are also used to create virtual desktops

Automating:

- Automating tasks :
 - provisioning, configuration, patching, release management and compliance.
- Automating tasks make data centers run more efficiently.

Securing:

- Security of physical infrastructures.
- Physical security, Network security, and Data and User security

Data Center Condition

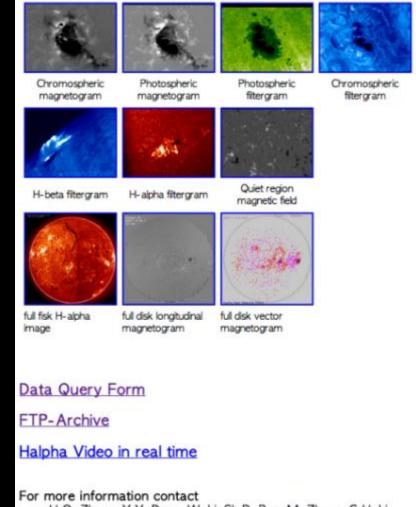
Storage

- Easy to Data Select and Translate
- Perfect information
- Pree from any jamming environments
- So-work between similar research group synergy effect
- ⁹ Contribute the research fields

http://www.ips.gov.au/World_Data_Centre



http://sun.bao.ac.cn/observation/obsdata_e.html



H.Q. Zhang, Y.Y. Deng, W. Li, Sh.D. Bao, M. Zhang, G.H. Lin

http://magdas2.serc.kyushu-u.ac.jp/datausage/index.html



Requesting Data and Rules for Usage

How To Get Data from SERC

We define some terms as follows: First Party...Space Environment Research Center, Kyushu University, Japan. Second Party...A location that is hosting MAGDAS-II instrument. Third Party...All others.

Second Party has the privilege to access 1-second realtime or corrected data of its station or stations.

Second Party also has the privilege to access 1-minute realtime data of its station or stations.

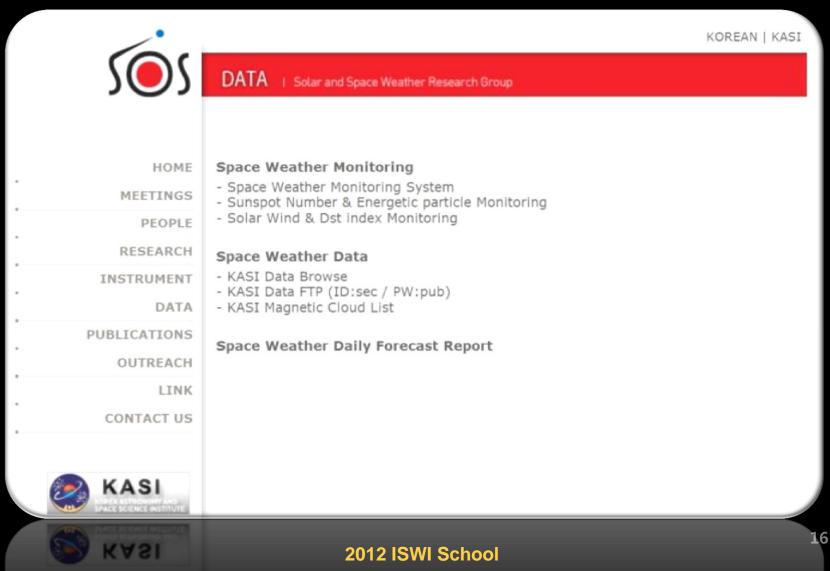
And Second Party has the privilege to access 1-minute corrected data of all the stations of the First Party. Data can be accessed via the ftp site of the First Party.

And Second Party has the privilege to access 1-minute corrected data of all the stations of the First Party. Qata can be accessed via the ftp site of the First Party.

http://magdas2.serc.kyushu-u.ac.jp/datausage/index.html

*********************** Data Request Form *	*****
TO THE ATTENTION OF PROF. K. YUMOTO . Professor of Kyushu University, Japan. 2. PI of the MAGDAS Project 3. Director of SERC	
Please print very clearly:	
Your Name:	
Your Position:	
Your Institute:	
Your telephone or fax number: please include country code)	
Your email address:	
Purpose of Request Data:	
	_

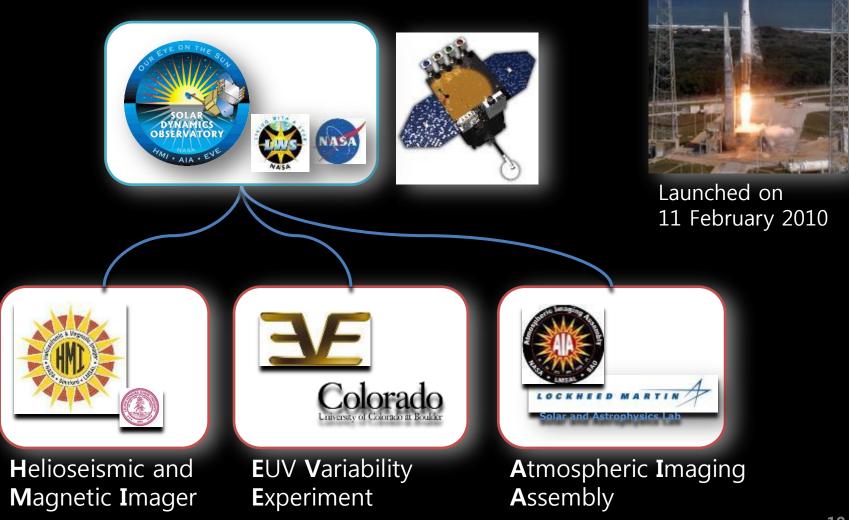
http://sos.kasi.re.kr/english/data.php



SDO and RBSP for SWx

SDO and RBSP

• What is SDO



SDO and RBSP Science

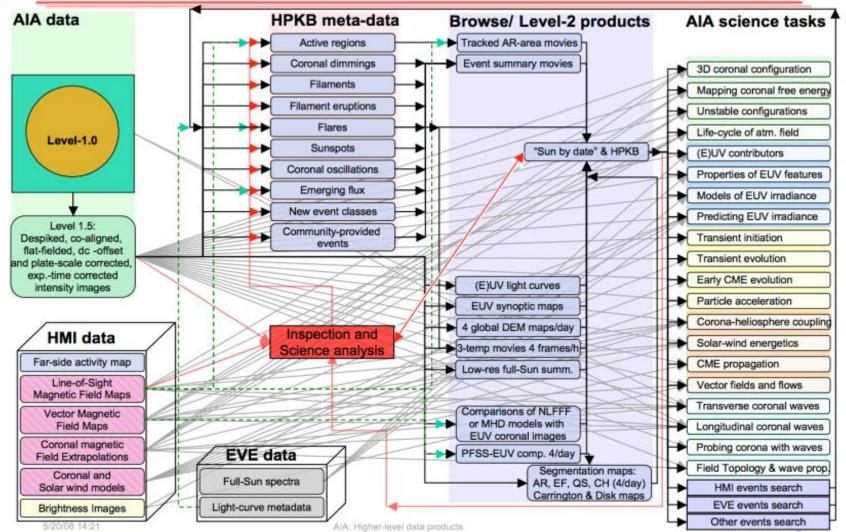
Mission Science Objectives of SDO

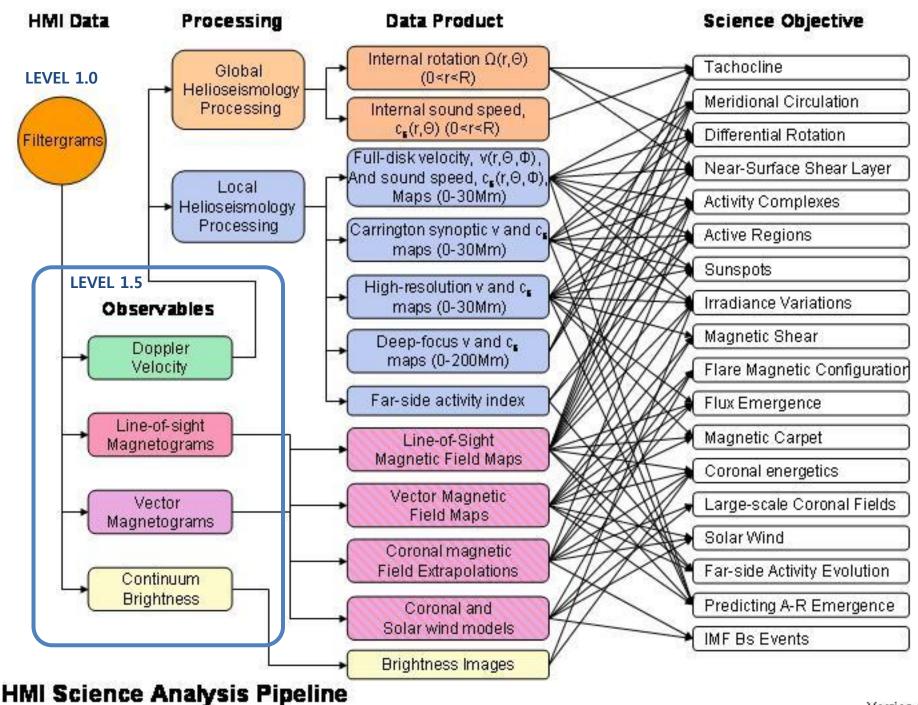
- 1. Solve the mechanisms of 11-year cycle of solar activity
- 2. How is active region magnetic flux synthesized, concentrated, and dispersed across the solar surface?
- 3. Magnetic reconnection on small scales
 - ♦ How significant is it in coronal heating
 - \diamond Aaccelerating the solar wind
- 4. Where do the observed variations in the Sun's EUV spectral irradiance arise, and how do they relate to the magnetic activity cycles?
- 5. What magnetic field configurations lead to the CMEs, filament eruptions, and flares that produce energetic particles and radiation?
- 6. Can the structure and dynamics of the solar wind near Earth be determined from the magnetic field configuration and atmospheric structure near the solar surface?
- 7. Reliable forecasts of space weather and climate, activity occur.



AIA Processing Pipeline and Standard Products







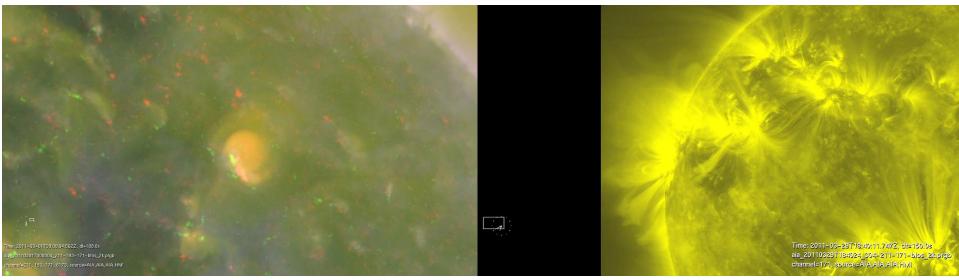
Version 1.0

1) AIA Flares (Period: 2010-04-24 ~ 2011-04-09)

Average time period per each flare: 10~20 min -> average 15 min
 ≥GOES C-class: 437 flare events
 ≥GOES M-class: 55 flare events
 ≥GOES X-class: 2 flare events

2) Eruption Events

- ♦ Event period : several hours (1-10 hours)



3) Coronal Jet :

- * 131 events
- Event period : ~ 1hr : 211, 171, 304 A

4) Spray Surge

- ✤ 85 events, Event period: ~ 1hr
- Typical wavelength : 304 A, multi wav: 211, 171

5) Filament

- ✤ 1 event/day, Event period: 24 hr
- * Main wavelength: 304 A, multi wav: 211, 171, 193 A

6) Filament Eruption

- ✤ 361 events : 1 event/day, Event period: ~5 hr
- ✤ Main wavelength: 304A, multi wave: 211, 193, 171 A

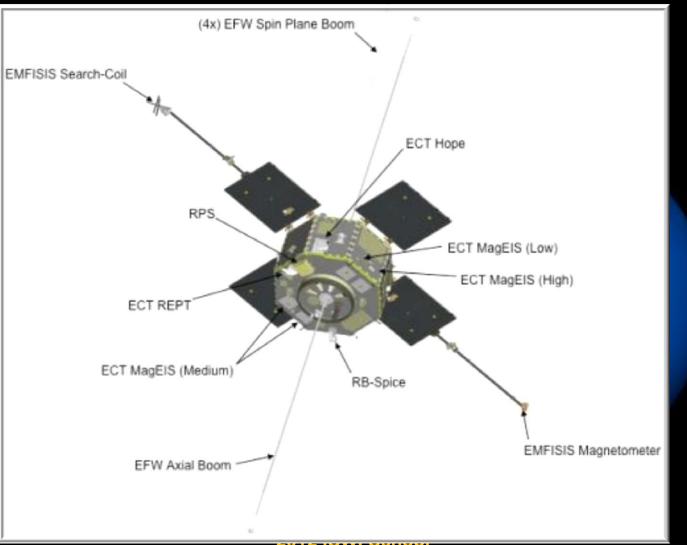
7) Filament Activation

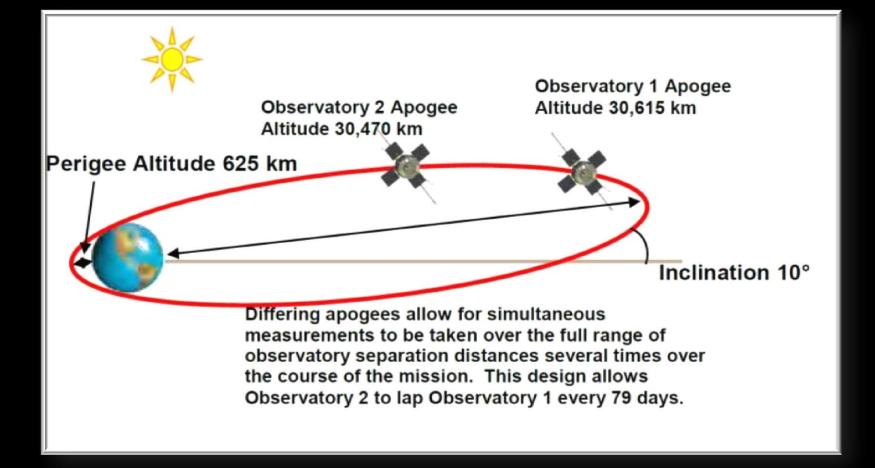
- * 120 events : 10 days per 1 month, Event period : 5 hours
- ✤ Main wavelength: 304 A, multi wavelength : 211, 193(304), 171 A

```
    ♦ Eye capture ~ 0.1% of real activity
    ♦ Eye estimate : 0.5Gb/day
```



RBSP Mission





- Spin stabilization (5 rpm)
 - spin axes pointing sun
- Orbit : 620 ~ 30,500 km
 inclination: 10°
 - (THEMIS: 1.5 x 10 Re)

Period: 9h

- Time scale comparable to

the storm development

RBSP payloads

Investigation	Instruments	PI
Energetic Particle Composition and Thermal Plasma Suite (ECT)	Helium Oxygen Proton Electron Spectrometer(HOPE) Magnetic Electron Ion Spectrometer (MagEIS) Relativistic Electron Proton Telescope (REPT)	H. Spence UNH
Electric and Magnetic Field Instrument Suite and Integrated Science (EMFISIS)	Low-Frequency Magnetometer (MAG) High-Frequency Magnetometer and Waveform Receiver(Waves)	C. Kletzing University of Iowa
Electric Field and Waves Instrument for the NASA RBSP Mission (EFW)	Electric Field and Waves Instrument for the NASA RBSP Mission(EFW)	J. Wygant University of Minnesota
Radiation Belt Storm Probes Ion Composition Experiment (RBSPICE)	Radiation Belt Storm Probes Ion Composition Experiment (RBSPICE)	L. Lanzerotti New Jersey Institute of Technology
Proton Spectrometer Belt Research (PSBR)	Relativistic Proton Spectrometer (RPS)	D. Byers NRO

The broad objective of RBSP

- Which Physical Processes Produce Radiation Belt Enhancement Events?
- What Are the Dominant Mechanisms for Relativistic Electron Loss?
- How do Ring Current and other geomagnetic processes affect Radiation Belt Behavior?

Application of RBSP data

0 Satellite Operation

- ~100 keV electrons are source of spacecraft anomaly (Choi et al. 2011).
- NOAA-GOES satellite is not useful.
- **Aviation** 0
 - Warning radiation risk
- Radio communication 0
 - Monitoring ring current will provide information for ionosphere disturbance.





SDO & RBSP Data Center in KASI

Why should KASI construct KDC for SDO?

9 For solar and space weather research

- High cadence and high resolution data
- Combine to FISS and HINODE

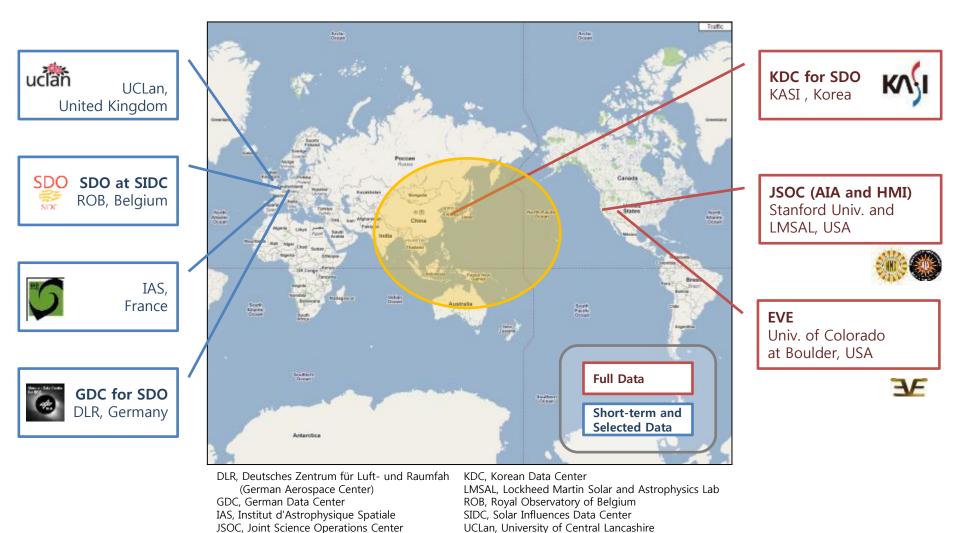
Yery large data size

- It is not easy to successfully download large scientific data of SDO from NASA.
- It is totally impossible through public network.

9 For space weather applications

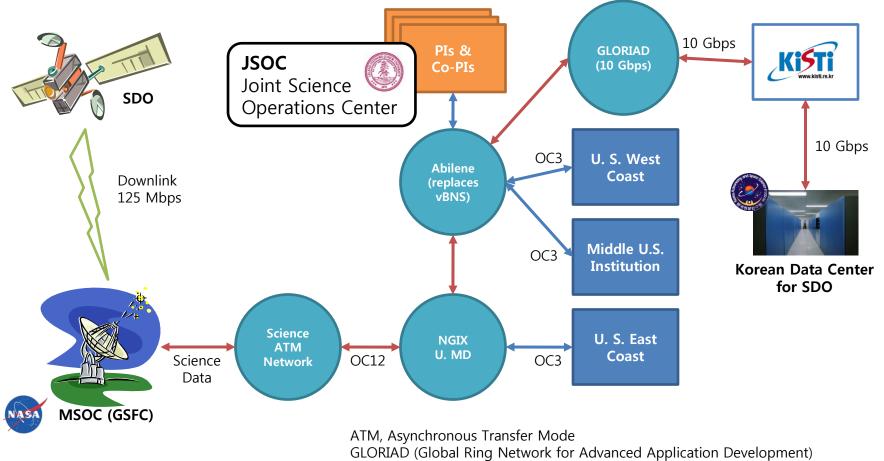
Real time or quasi real time SDO data

Data Centers for SDO



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NASA-JSOC-KASI Network



MSOC, Mission and Science Operations Center

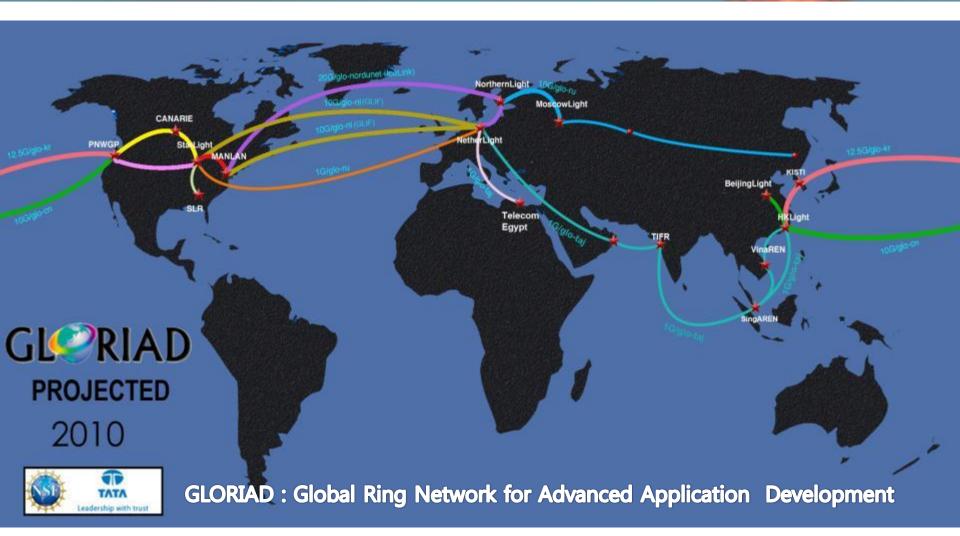
NGIX, Next Generation Internet eXchange

OC, Optical Carrier (OC3: 155 Mbps, OC12: 622 Mbps)

U. MD, University of Maryland

vBNS, very-high-performance Backbone Network System

GLORIAD



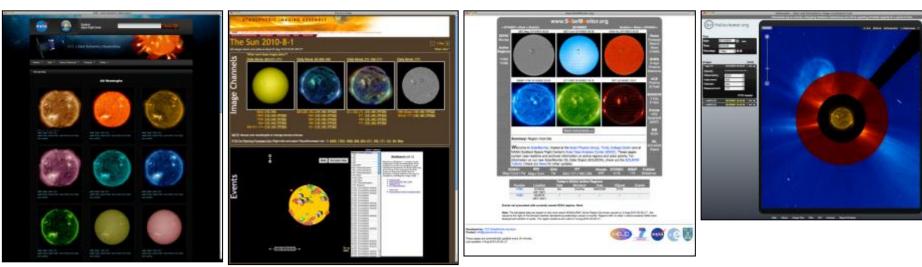
SDO Data Browse

The Sun Now

The Sun Today

Solar Monitor

Helioviewer



- The Sun Now
- Solar Monitor
- Helioviewer
- JHelioviewer

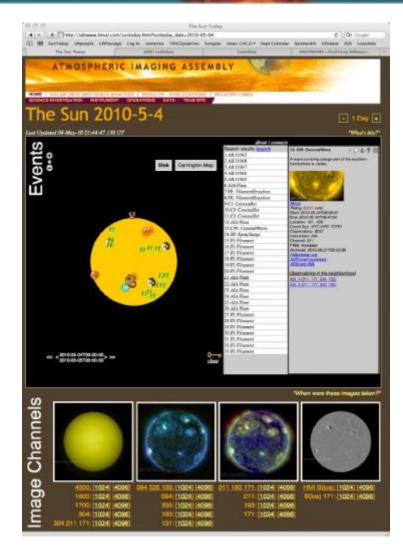
- : http://sdo.gsfc.nasa.gov/data/
- The Sun Today : http://sdowww.lmsal.com/
 - : http://solarmonitor.org/
 - : http://helioviewer.org/
 - : http://jhelioviewer.org/

JHelioviewer



iSolSearch, LMSAL

- HEK (Heliophysics Events Knowledgebase)
- Daily summary of the solar state
- Events and features
- Multi-thermal images
- Links to detailed descriptions
- Firefox 3.0+, Safari 3.2+, Chrome



FESTIVAL



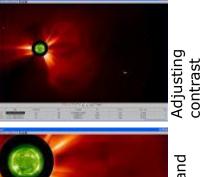
http://www.ias.u-psud.fr/stereo/festival/

- IDL-based browser
- Open source
- Dynamic composite images
- Solar Data : SECCHI/STEREO, EIT/SOHO, LASCO/SOHO, NRH, MkIV, TRACE, XRT
- Output : PS, JPEG, MPEG, AVI, etc

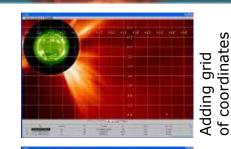
The upgrade version will support SDO .



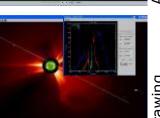
Combining solar data



Zooming and panning







Drawing a profile

mask filters

Running difference

How to Get AIA and HMI Data

Lookdata Tool

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? You may go directly to 1	Step 3 on the above Records	et Select tab if you know which series you want.	Itter Date 2
1. Find list of dataseries		2. Pick series to use	Batter 19
Enter a dataseries match pattern to search for seriestnesses, or leave black to select from all series.		Seloct data series here.	
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Cutout Service

os Coverage Registry (HCR 600 Status: The ASA and HHS data are not yet fully calibrated, but test series are vallable for scientists to see the headers and otherwise toot their compatibility with their tools. We have not yet started on EVE integration leastly which ready for writer Cart 30. 160 Married Wesley House, Search for Solar Physics Data Products: i pakes and Sch If pay're new in the VEO, see mow To Saarch, the FAQ or click the ¹⁸ issues for onlin 10-19-19 ST Please select which values you wish to use to search for data products 44.54 Time Search by time intervel 10.10.10 Derive time tritercals from event satalogs Observable Search based on physical observables® Destrument / Rearce / Provider AN COM Search based on instruments[®] or data antives[®] 386 C1968 C1970 C269 C299 M(0) C298 C294 M(04 C2600 Instrument / Source (not provider dependent) a Instrument Only (net source or provider dependent) Spectral Range Search based on a spectral range IN CONT Nicksames Search based on common terms used to describe data products. Note: Nicknames generate an intersection with other search terms, so searching for a nickname, and a physical siz for other parametert when a nickname defines other physical observables will result in no matches Show Mickname Definitions Searching against current VSD instance Income will insuit from 1 **VSO Documentation** Documentation for locaritety, Programmers and Data Providers, Including Changes, FADs, and contact info Help us improve VSO · Tell us what features you would like to see. Other suggestions / comments / criticiam Cantact information for VSD team members clear toout-date that that date Inter-36 YOeWar-ROT VSD B Home | NSO | Stanford on Different Level 1.5 data is served from MW Level 1.5 Advance from 2000 WIT THE Automatically Generated at 1 Bled Dol 27 81:17-07 2020

Virtual Solar Observatory

- Lookdata Tool : http://jsoc.stanford.edu/ajax/lookdata.html

Rowald (1) Roward Tille (1) Max Premes (1)

- Cutout Service : LMSAL
- VSO : http://vso.nso.edu/

KASI SDO Website(sdo.kasi.re.kr)



KDC for SDO

After arrival, the spacecraft spent several months using a technique called aerobraking, which involved dipping into the Martian atmosphere to adjust its orbit. In February 2002, science operations began.

Data

- . Discovery sessions during the SVA encompassed multiple teams at Redbox.
- # 4096 PFSS 2048 PFSS 1024 PFSS 512 PFSS 48 hr MPEG
- Commercial Spaceflight 60 Day Report, June 29, 2011
- EUVI / AIA 195 Stonyhurst Heliographic
- Commercial Market Assessment (CMA) for Crew and Cargo Systems, April 2011 (742K PDF)

Browse Data 🔘

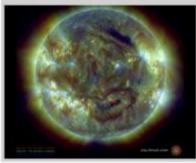






Courtesy of NASA/SDD and the AIA, EVE, and HM science teament. Korean Data Center for SDD, Korea astronomy & space science institute 776 Daedeokdae-ro, Yuseong-gu, Daejnon, 305-348. Rep. of Korea

The Sun Now



KASI SDO Website(sdo.kasi.re.kr)



Dates

2012-09-05 01:00:00

2012-09-05 02:00:00

Available from 2012-07-01

Telescopes

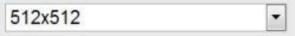
AIA 171 (gold)

Choose a display

Images O Movie Archive (FITS/JPG)

•

Resolution



Submit 📀

2012 ISWI School

KASI SDO Website

- 1. Based on NASA SDO + LMSAL web-design
- 2. Provide same type of Lookdata Tool in Stanford
- 3. Provide solar viewer S/W like Jhelioviewer
- 4. Large amount library
- 5. Connect to VSO
- 6. Open : Jan. 1st 2013 http://sdo.kasi.re.kr

Research power of KASI

- 9 # of group members : 26
- 9 71 SCI papers by 7 scientist during last 3 yrs. from 2009 to 2011
- Solar physics (30), Space science(41)
- Strong part : CME, solar flare, ionosphere.
- Weak point : space weather model, MHD simulation in solar physics.
- > Scientific Joint Group by using SDO and RBSP data
- **WE ARE WELCOME TO CO-WORK PARTNER IN ANYTIME**
- http://sos.kasi.re.kr/



2012 ISWI School

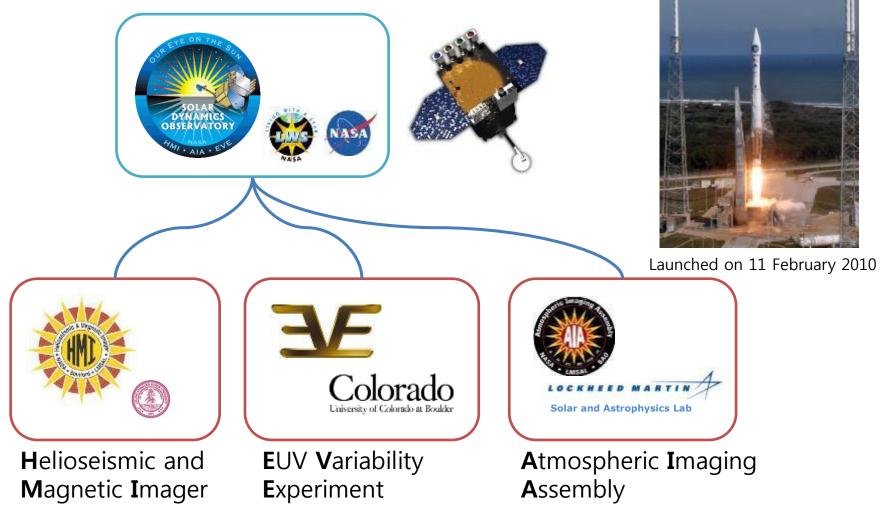
RBSP First Science Endeavors

What issues can be resolved about strong and weak whistler mode interactions and their roles in electron energization and loss in the first 3 months?

What issues can be resolved about the large scale dynamics and structure with just the first few major geomagnetic storms?

What issues can be resolved about the source, structure, and dynamics of the inner (L<2) ion and electron belts in the first 3 months?

SDO, Solar Dynamics Observatory



SDO Data Size

Instruments	Mbps	MB/day
HMI	25	270,000
AIA	50	540,000
EVE SIM	1	10,800
Coronagraph	1	10,800
UV/EUV Spectrometer	30	324,000
Photometric Mapper	5	54,000
Vector Magnetograph	5	54,000
Total	117	1,263,600

AIA Data Volume

TRACE Data Volume
EIT Data Volume

HMI Data Volume

X Source : SDO Science Definition Team Report, 43 page

- One Day : 1.2 TB/day
- One Year : 438 TB/year
- Five Years : 2.1 PB

Data Products from AIA

- Pata at Level 0 are images that have been constructed from the raw telemetry stream.
- Pata at Level 1.0 are images that have been converted from Level 0, with processing including bad-pixel removal, despiking and flat-fielding. All higher-level products for AIA are based on Level 1.0 data.
- Pata at Level 1.5 are images that have been adjusted to a common 0.6" plate scale, and that share common centers and rotation angles, but are not exposure-time corrected

Data Products from HMI

- Pata at Level 0 are images that have been constructed from the raw telemetry stream.
- Pata at Level 1.0 are images that have been converted from Level 0, with processing including bad-pixel removal, flat-fielding, and quality assessment checks, but otherwise not having undergone any irreversible data alterations.
- Pota at Level 1.5 are images of the physical observables (Dopplergrams, magnetograms, and continuum images), which were constructed from the individual Level 1.0 filtergrams.
- Pata at Level 2 have been irrevocably filtered, time-sequence-merged, Fourier-transformed or otherwise changed from Level 1.5 data in a way that is irreversible. Level 2 products include intermediate products for later production of mission science data products, such as helioseismic inferences of solar subsurface flows.

Korea-Japan Network

